



## Complexity, turbulence and stochasticity in geospace plasma environment: role on predictability

Giuseppe Consolini<sup>(1)</sup>

(1) INAF-Istituto di Astrofisica e Planetologia Spaziali, Roma, Italy

The geospace plasma environment is a nonequilibrium system characterized by a complex dynamics in response to the variations of the interplanetary conditions (solar wind and interplanetary magnetic field) [1, 2, 3]. This complex character manifests in turbulent, stochastic and near criticality fluctuations of magnetic field and plasma parameters in almost all the magnetospheric and ionospheric regions, as well as, in proxies of the geomagnetic activity (e.g., geomagnetic indices). For instance, in 2017 Alberti et al. [4] evidenced that the fluctuations at distinct timescales of some geomagnetic indices differently respond to the changes of the interplanetary conditions during magnetic storms, being representative of different processes occurring in the plasma geospace. Now, one of the main features of these fluctuations is their multiscale and self-similar nature that in many different situations may strongly influence the homogeneity of the plasma environment generating multiscale plasma structures, whose dynamics acts on the evolution of the overall geospace plasma environment especially during the occurrence of magnetic storms and magnetospheric substorms.

The understanding and modelling of the multiscale and complex character of the magnetosphere–ionosphere (MI) dynamics is crucial for developing reliable models capable of correctly forecasting the effects of a solar disturbance [5]. Indeed, the dynamical complexity of the geospace plasma environment has a number of crucial effects on the predictability of the magnetosphere–ionosphere, limiting *de-facto* the capability of forecasting the geospace plasma status at certain spatio-temporal scales.

Here, we present an overview on the complex dynamics of the geospace plasma environment as monitored by a set of geomagnetic indices, and *in-situ* measurements, and discuss the impact of the multiscale character of the magnetosphere–ionosphere dynamics on limiting the forecast horizon of the geospace plasma dynamics. The latter point is examined by applying concepts from the theory of dynamical systems, such as the correlation dimension  $D_2$ , the Kolmogorov entropy  $K_2$  [6], etc.

### References

- [1] G. Consolini, P. De Michelis and R. Tozzi, “On the Earth’s magnetospheric dynamics: Nonequilibrium evolution and the fluctuation theorem”, *Journal of Geophysical Research: Space Physics*, **113**, A8, August 2008, CiteID A08222, doi: 10.1029/2008JA013074
- [2] G. Consolini, M.F. Marcucci and M. Candidi, “Multifractal Structure of Auroral Electrojet Index Data”, *Physical Review Letters*, **76**, 21, May 1996, pp. 4082–4085, doi: 10.1103/PhysRevLett.76.4082
- [3] G. Consolini and T.S. Chang, “Magnetic field topology and criticality in geotail dynamics: relevance to substorm phenomena”, *Space Science Reviews*, **95**, 1/2, January 2001, pp. 309–321, doi: 10.1023/A:1005252807049
- [4] T. Alberti, G. Consolini, F. Lepreti, M. Laurenza, A. Vecchio and V. Carbone, “Timescale separation in the solar wind–magnetosphere coupling during St. Patrick’s Day storms in 2013 and 2015”, *Journal of Geophysical Research: Space Physics*, **122**, 4, April 2017, pp. 4266–4283, doi: 10.1002/2016JA023175
- [5] T. Alberti, G. Consolini, P. De Michelis, M. Laurenza and M.F. Marcucci, “On fast and slow Earth’s magnetospheric dynamics during geomagnetic storms: a stochastic Langevin approach”, *Journal of Space Weather and Space Climate*, **8**, id.A56, December 2018, doi: 10.1051/swsc/2018039
- [6] G. Consolini, T. Alberti and P. De Michelis, “On the Forecast Horizon of Magnetospheric Dynamics: A Scale-to-Scale Approach”, *Journal of Geophysical Research: Space Physics*, **123**, 11, November 2018, pp. 9065–9077, doi: 10.1029/2018JA025952