



Global assimilative ionospheric modeling and the Tiny Ionospheric Photometer on the COSMIC constellation

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Abstract



- The Constellation Observing System for Meteorology, Ionosphere, and Climate (COSMIC) uses the Tiny Ionospheric Photometer (TIP) to characterize the nighttime ionosphere. TIP is a compact, nadir directed, narrow-band, ultraviolet photometer operating at the 135.6 nm wavelength. This emission is produced by recombination of O^+ ions and electrons, which is the natural decay process for the ionosphere. At night, the strength of the emission is proportional to the square of the peak electron density. TIP measures the horizontal structure of the ionosphere with 15-30 km resolution and high sensitivity, providing remarkable detail even during solar minimum conditions. With six sensors on COSMIC providing global coverage, TIP is a valuable resource for evaluating ionospheric models. We present a comparison of TIP data with models which employ global assimilation of ionospheric measurements. Ground-based GPS observations are routinely ingested and serve as the primary data source for assimilative models. With the launch of the COSMIC constellation in 2006, research has begun on assimilating GPS radio occultations. These space-based observations provide global coverage and promise to improve ionospheric specification in both a vertical and horizontal sense. Comparison with TIP reveals the horizontal performance of these models and demonstrates the improvement obtained from ingesting radio occultation data.



Overview



- *Does ingesting the Radio Occultation data improve the horizontal specification of the equatorial ionosphere?*
- FORMOSAT-3/COSMIC
 - Radio occultation (RO)
 - Tiny Ionospheric Photometer (TIP)
 - TIP serves as coincident, independent truth for evaluation of models assimilating RO data
 - Early in mission (best case)
 - 4 spacecraft are in same orbital plane
 - equivalent to constellation with 24 spacecraft
 - Post-sunset Equatorial Ionospheric Anomaly (difficult case)
- Comparisons of TIP with GAIMUSC
 - GAIM ingests COSMIC RO data and ground GPS data
 - IFM and IRI also included in analysis



FORMOSAT-3 COSMIC Overview

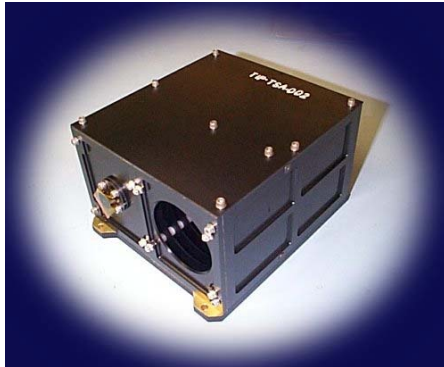


- **Constellation Observing System for Meteorology, Ionosphere and Climate**
- **April 2006 Launch**
 - Six spacecraft placed into initial 500 km altitude
- **Final Configuration as of December 2007**
 - Six orbit planes
 - 72 degree inclination
 - 800 km altitude
- **Sensors**
 - GPS Occultation
 - Tiny Ionospheric Photometer
 - Tri-Band Beacon
- **Satellite Operations Control Center in Taiwan**
- **Data Analysis and Archive Center in Boulder, Colorado**
- **3-year Mission (5-year Goal)**



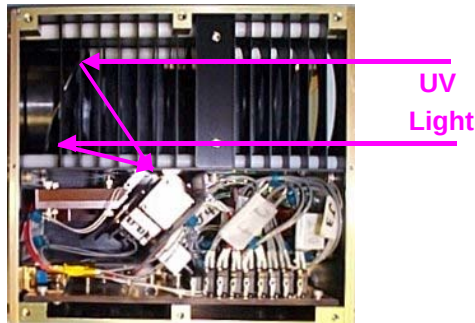


TIP Payload Overview



Sensor Module:

- Photomultiplier tube observes UV light
- Strontium fluoride filter passes 131-160 nm emissions
- Very high sensitivity ~600 counts/s/Rayleigh

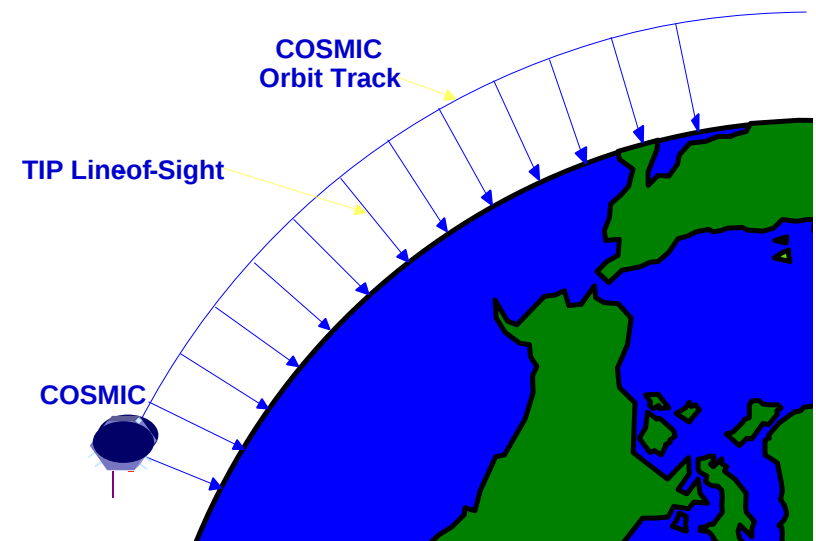


Measures intensity of naturally occurring 135.6-nm airglow, caused by the decay of the nighttime ionosphere.

Intensity is proportional to the vertical sum of the electron density-squared,

$$I \propto \int N_e^2$$

Nadir pointing 3.8° circular field-of-view yields 30-km horizontal resolution from an altitude of 800 km (15-km at 500-km altitude).

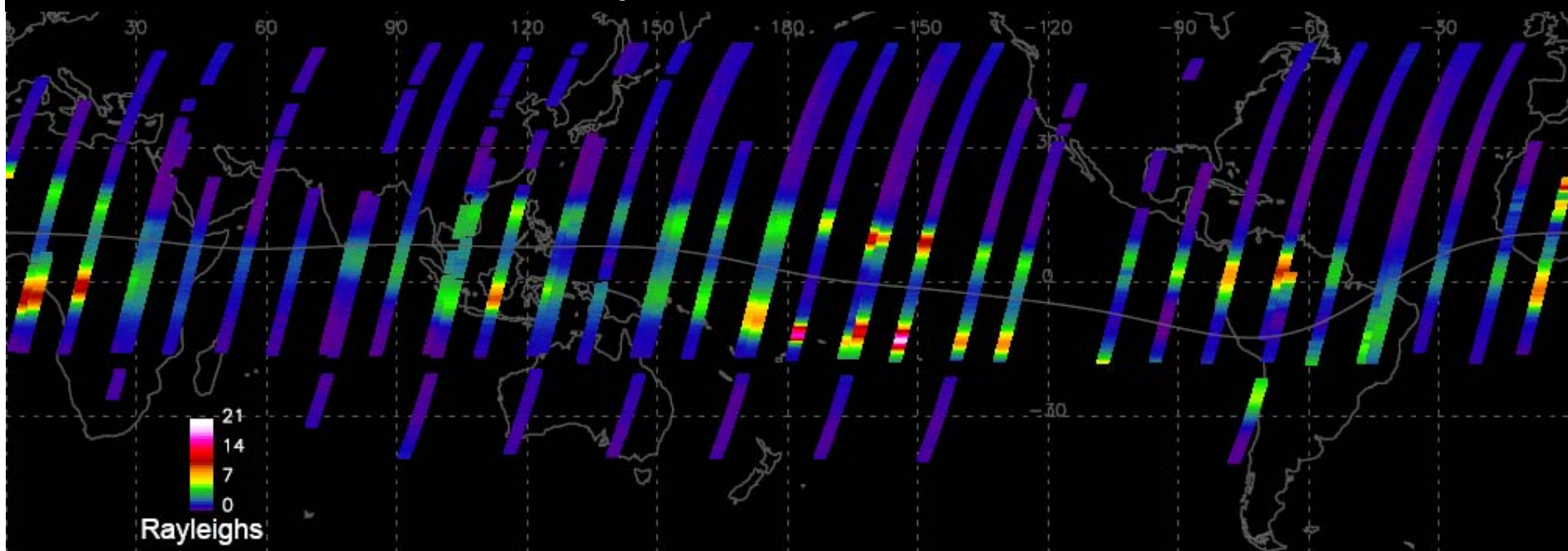




Variability of Post-Sunset Equatorial Ionosphere

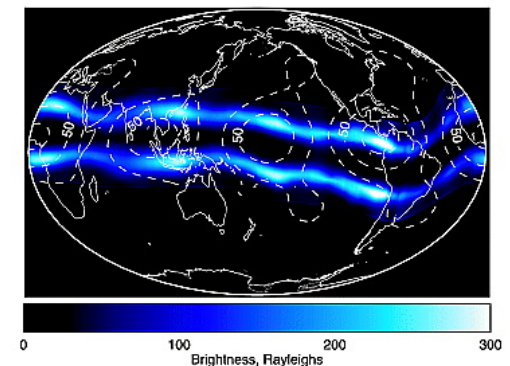


One-Day Composite FUV Image of Ionosphere at 2100 local time from multiple LEO spacecraft on COSMIC



TIP reveals the complexity of low latitude ionospheric drivers, vertical drift and meridional winds, as a function of longitude

Global 4-cell pattern is similar to observations by IMAGE FUV at equinox, solar maximum – attributed to atmospheric tides by Immel *et al.*, 2006



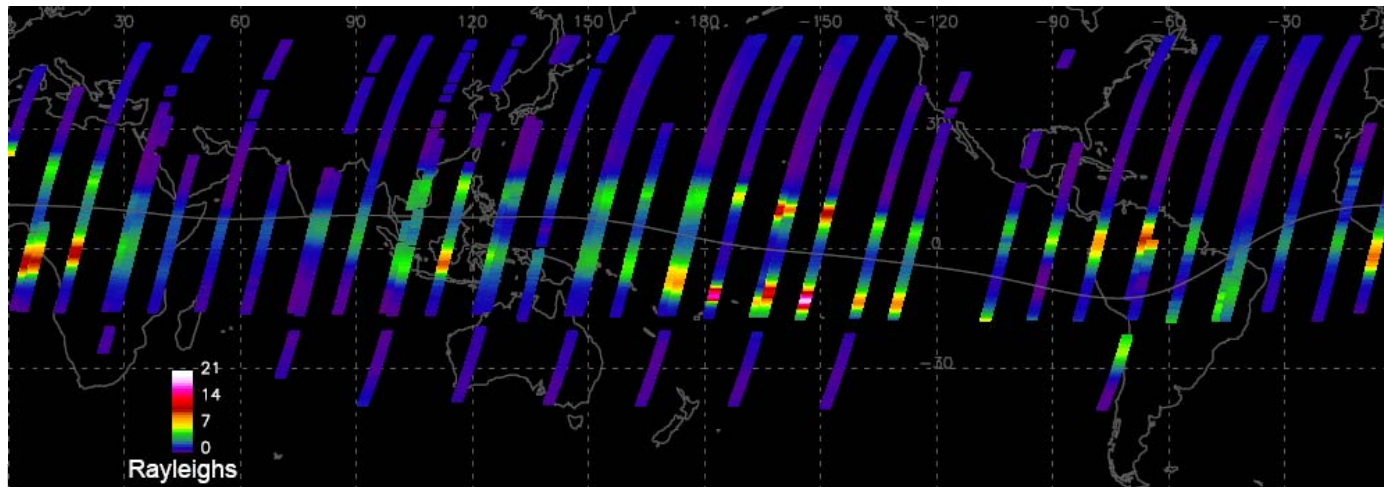


TIP and IFM Climatology

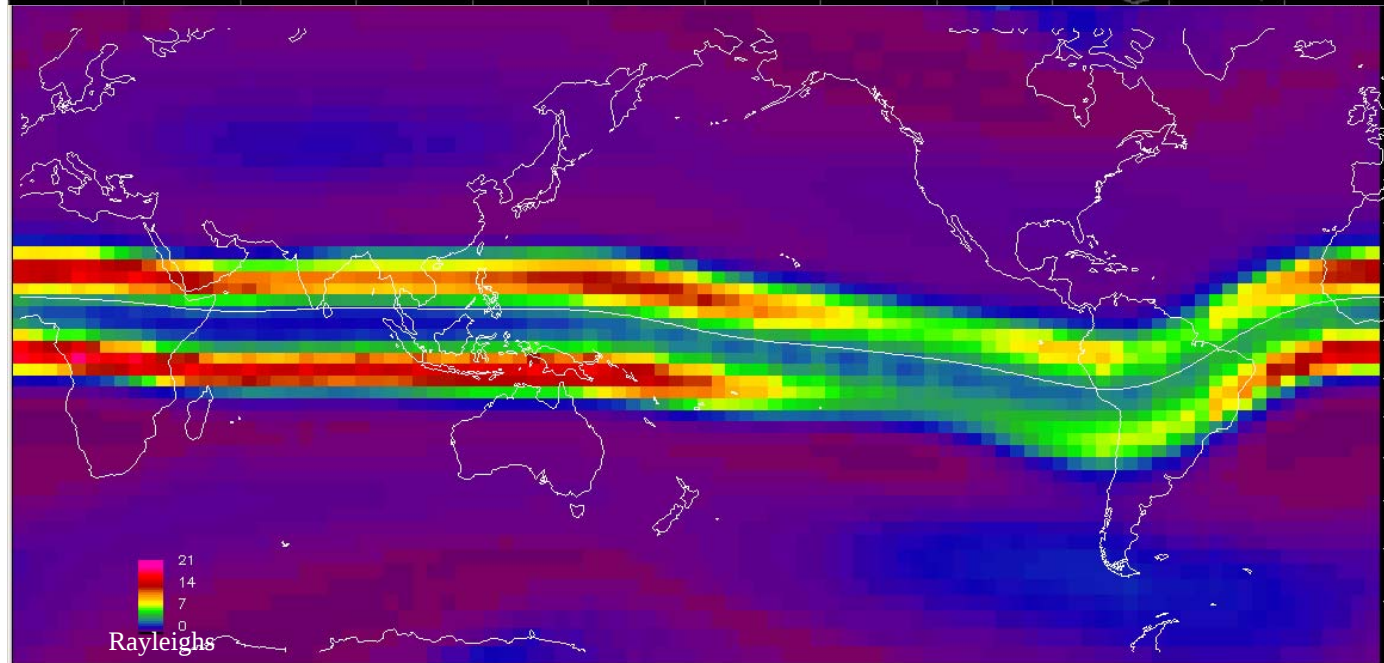
2100LT for 14 Sep 2006



135.6-nm
Radiance from
TIP



135.6-nm
Radiance from
IFM



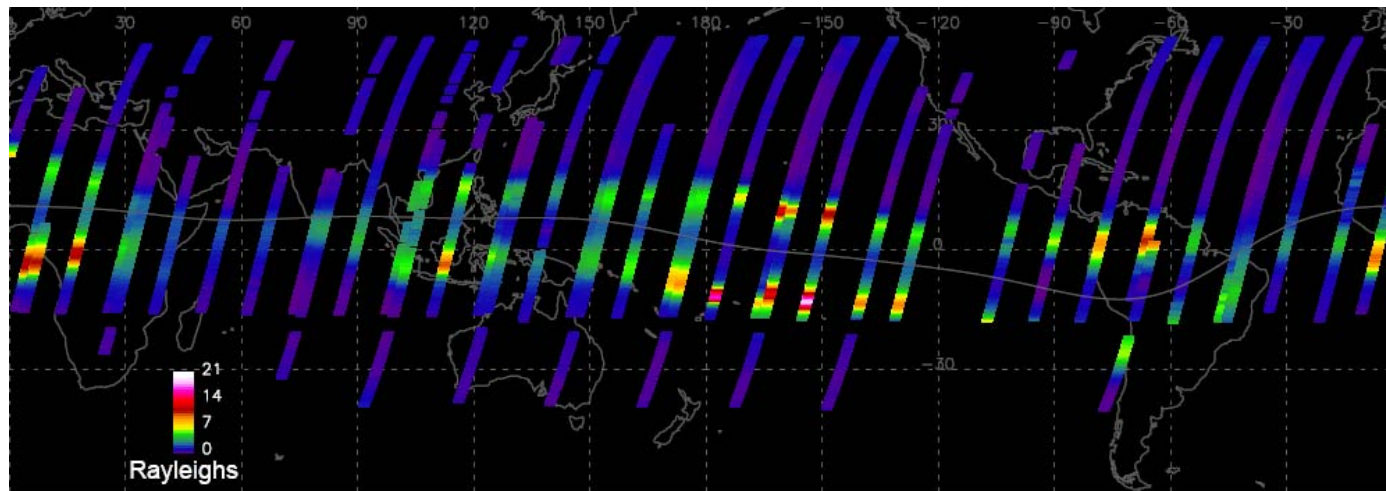


TIP and IRI Climatology

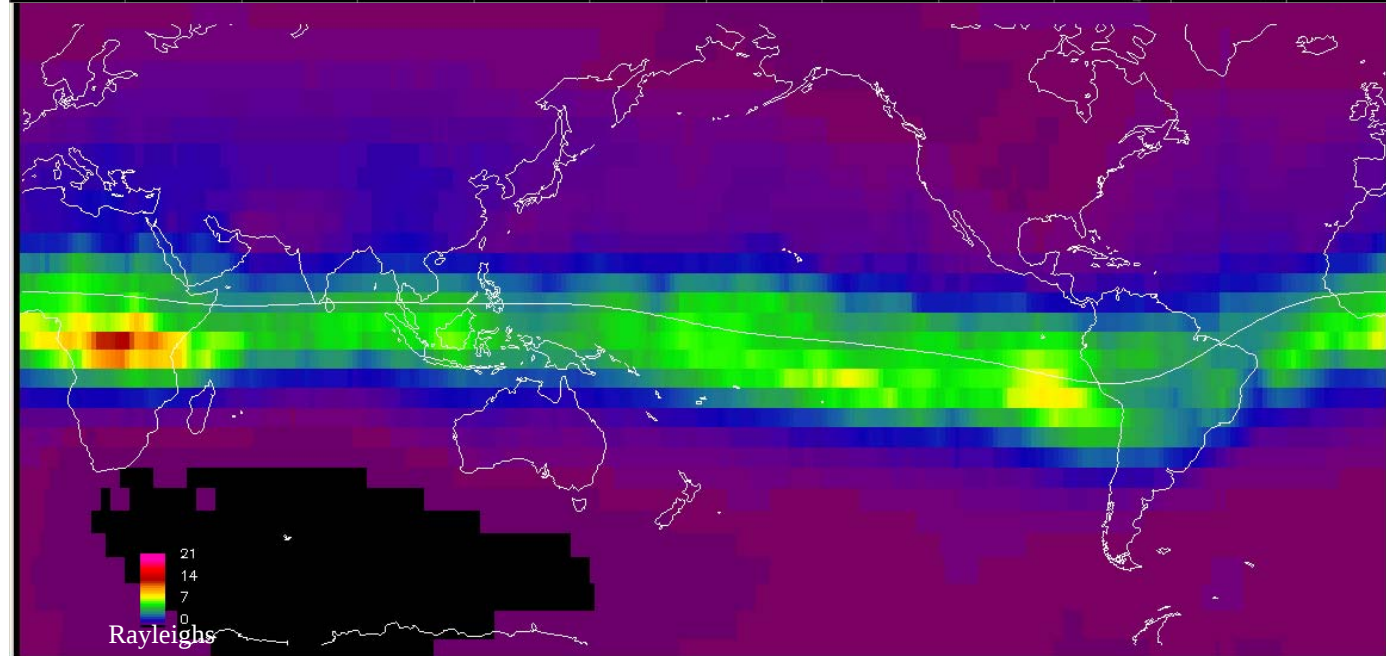
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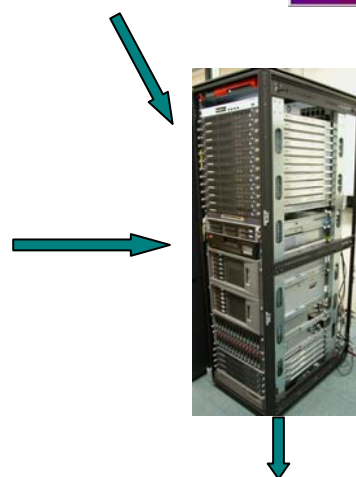
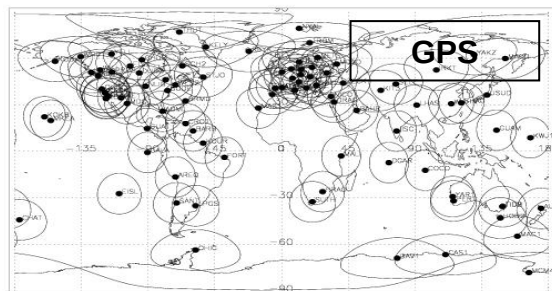
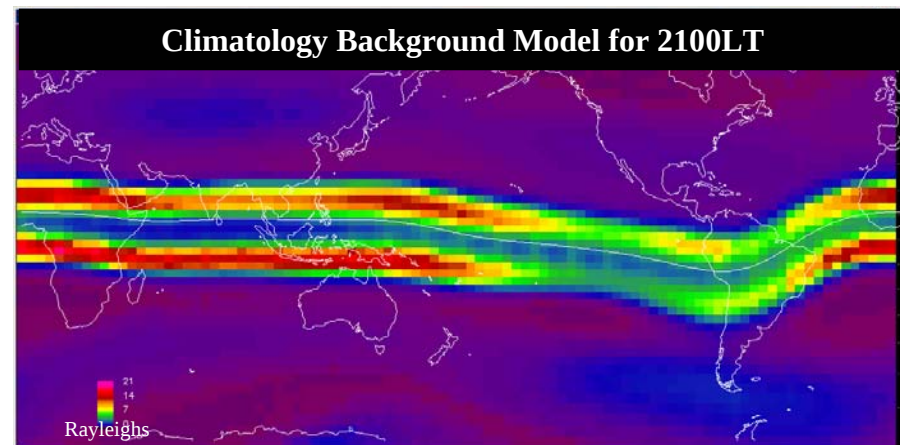
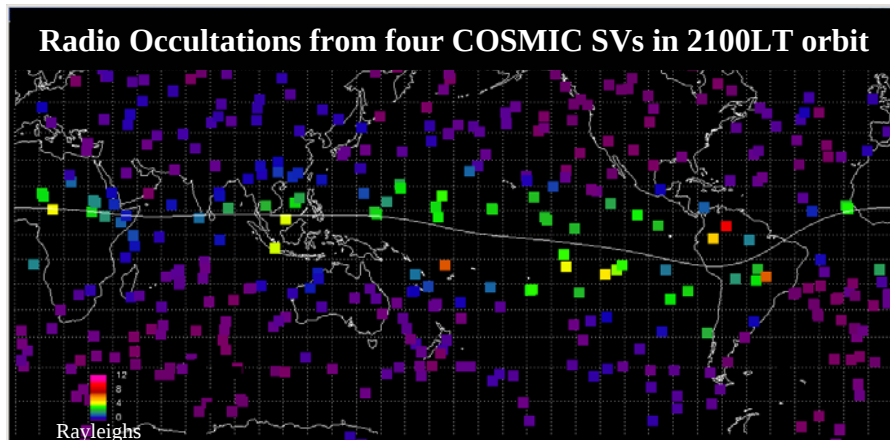


135.6-nm
Radiance from
TIP

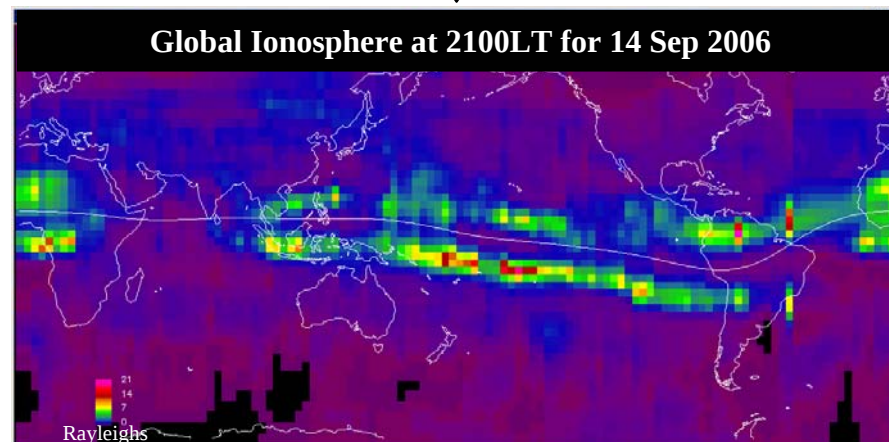


135.6-nm
Radiance from
IRI





GAIMUSC
Global Assimilation of
Ionospheric Measurements



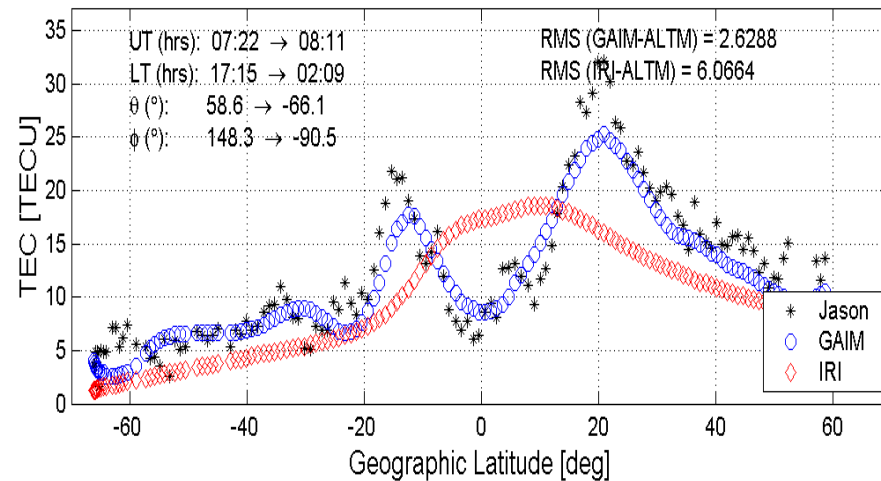


Validation of GAIM Using Altimeter TEC Measurements



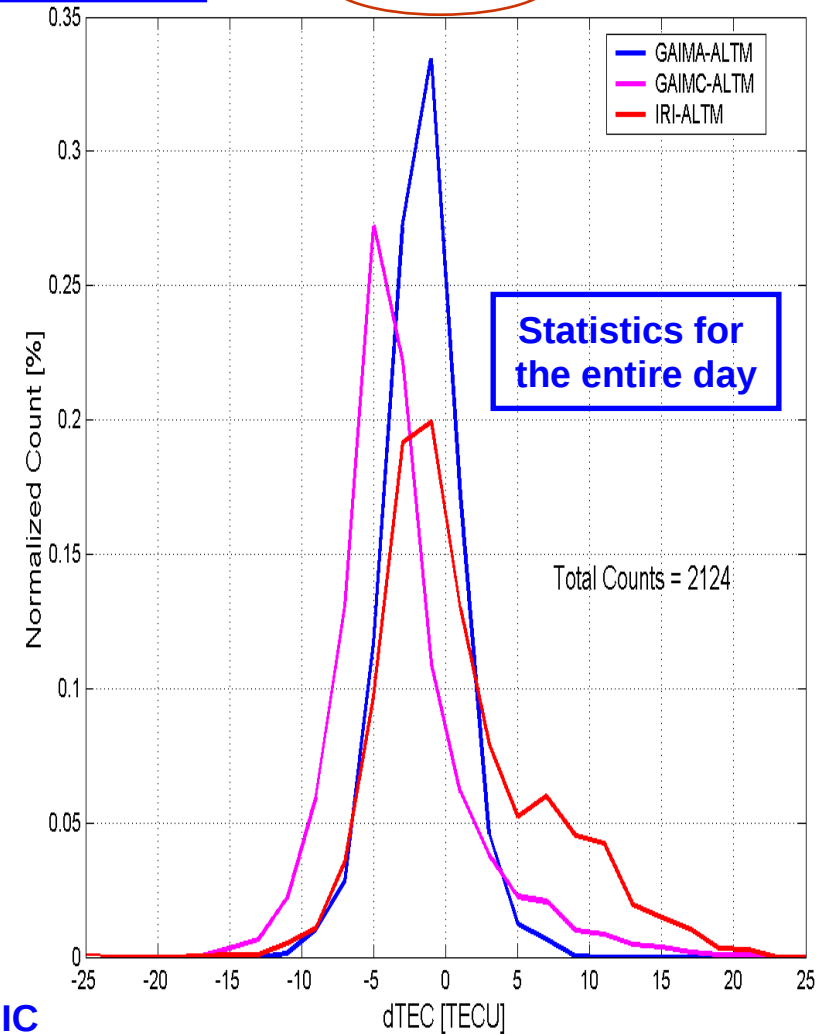
08/07/2006

TEC along Jason orbit tracks

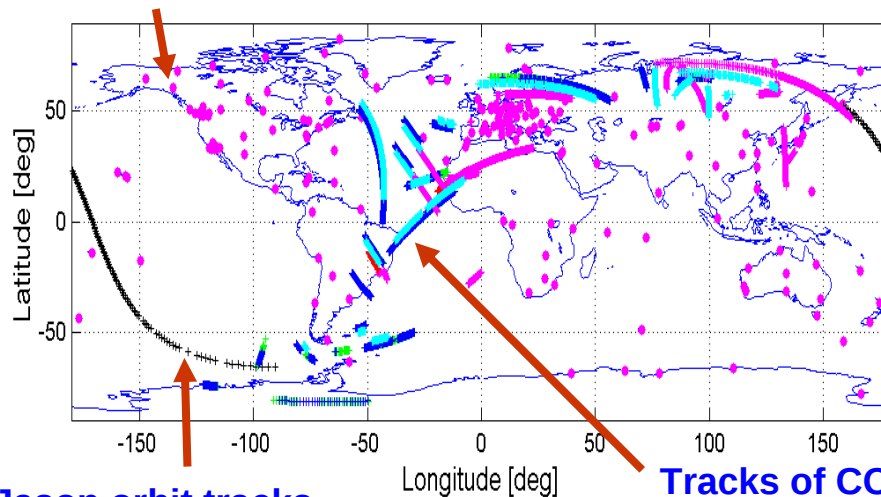


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Storm Day



GPS grnd station

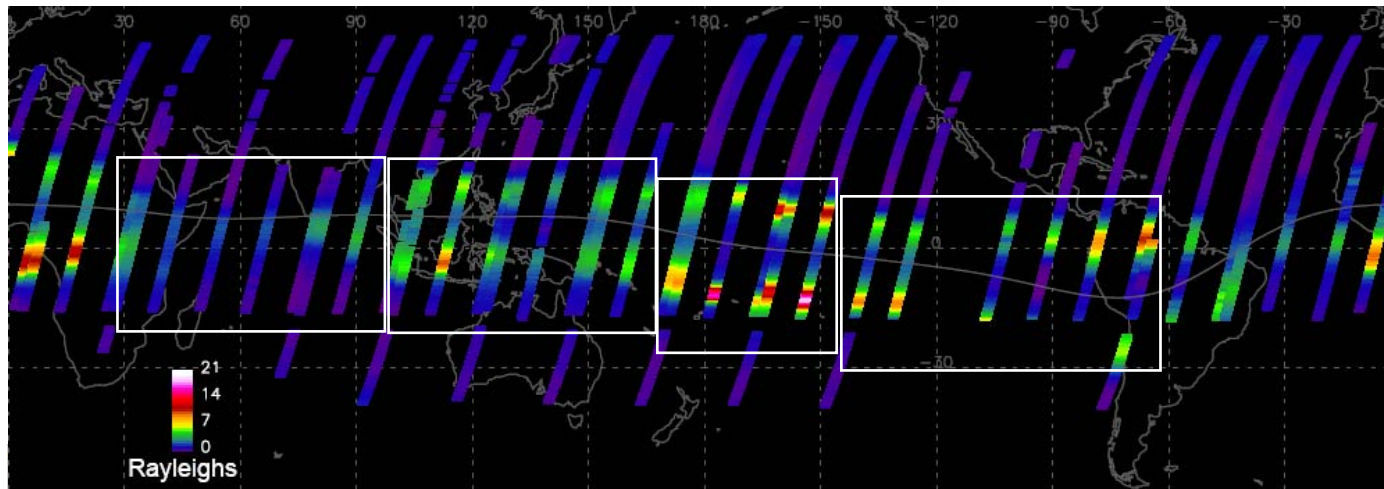




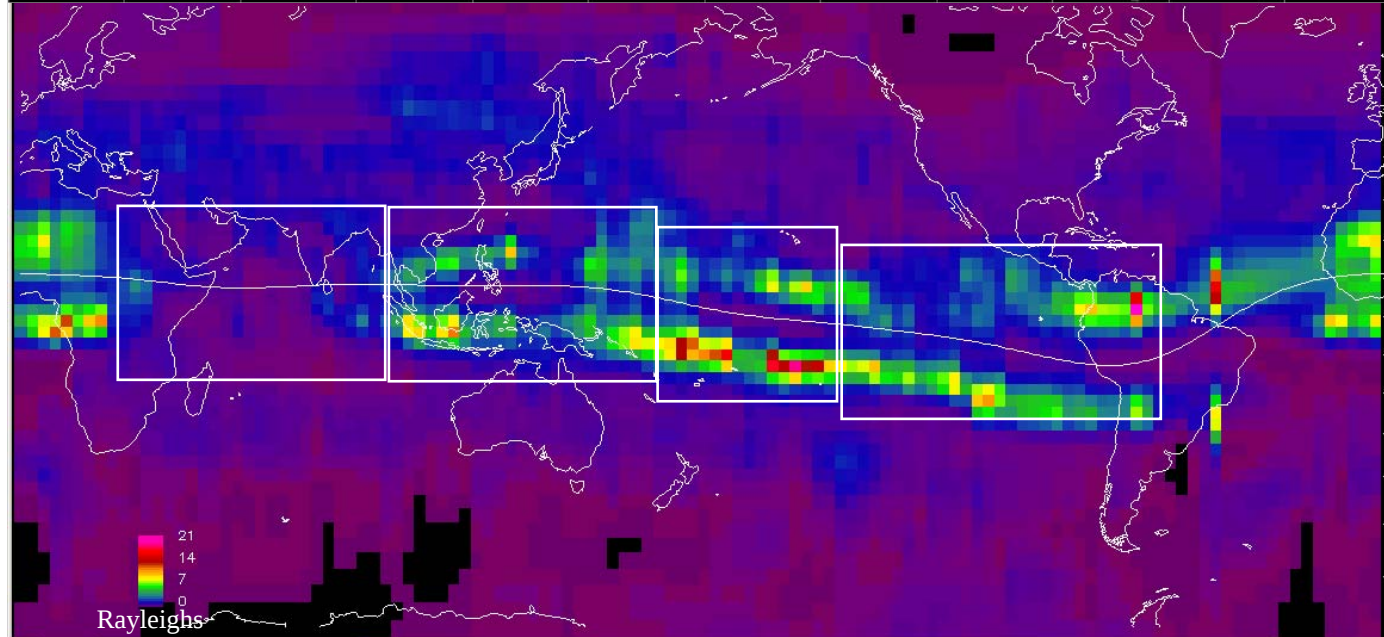
GAIMUSC Shows Remarkably Similar Non-migrating Tide Pattern on 14 Sep 2006



135.6-nm
Radiance from
TIP



135.6-nm
Radiance from
GAIM profiles



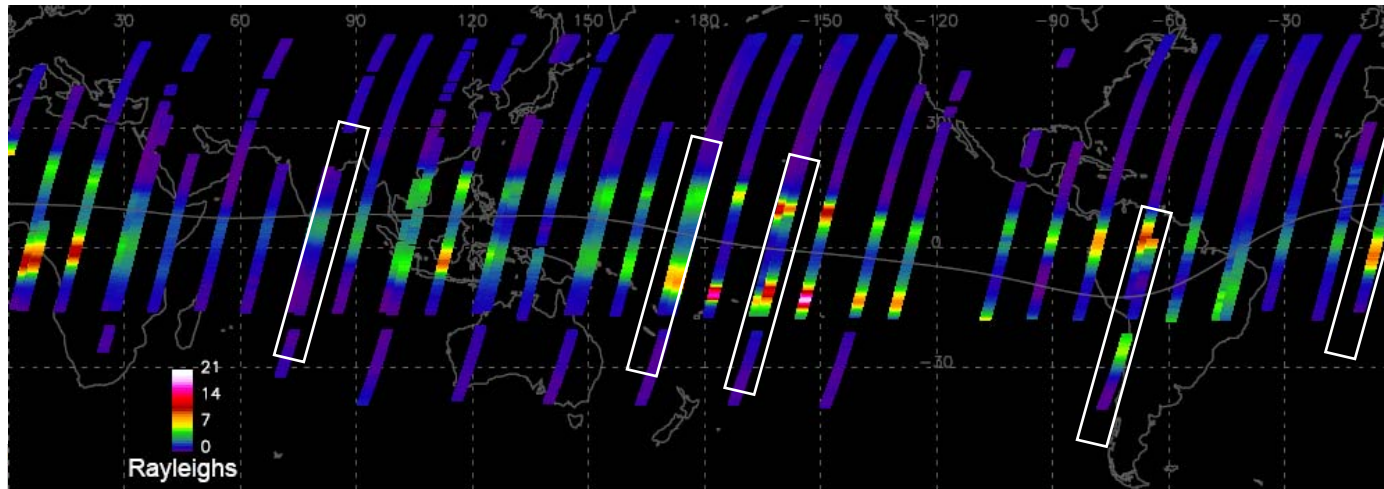


Comparison of TIP and GAIMUSC

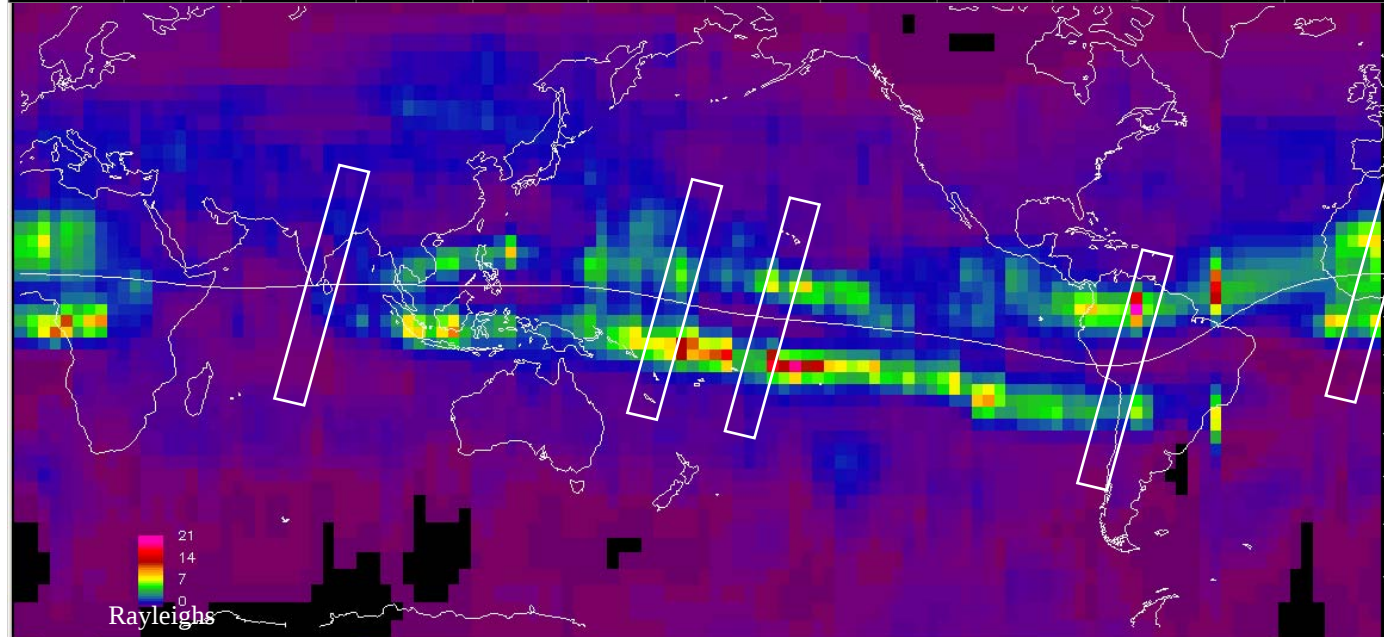
14 Sep 2006 FM1 FM3 FM6 (2100 LT)



135.6-nm
Radiance from
TIP



135.6-nm
Radiance from
GAIM profiles

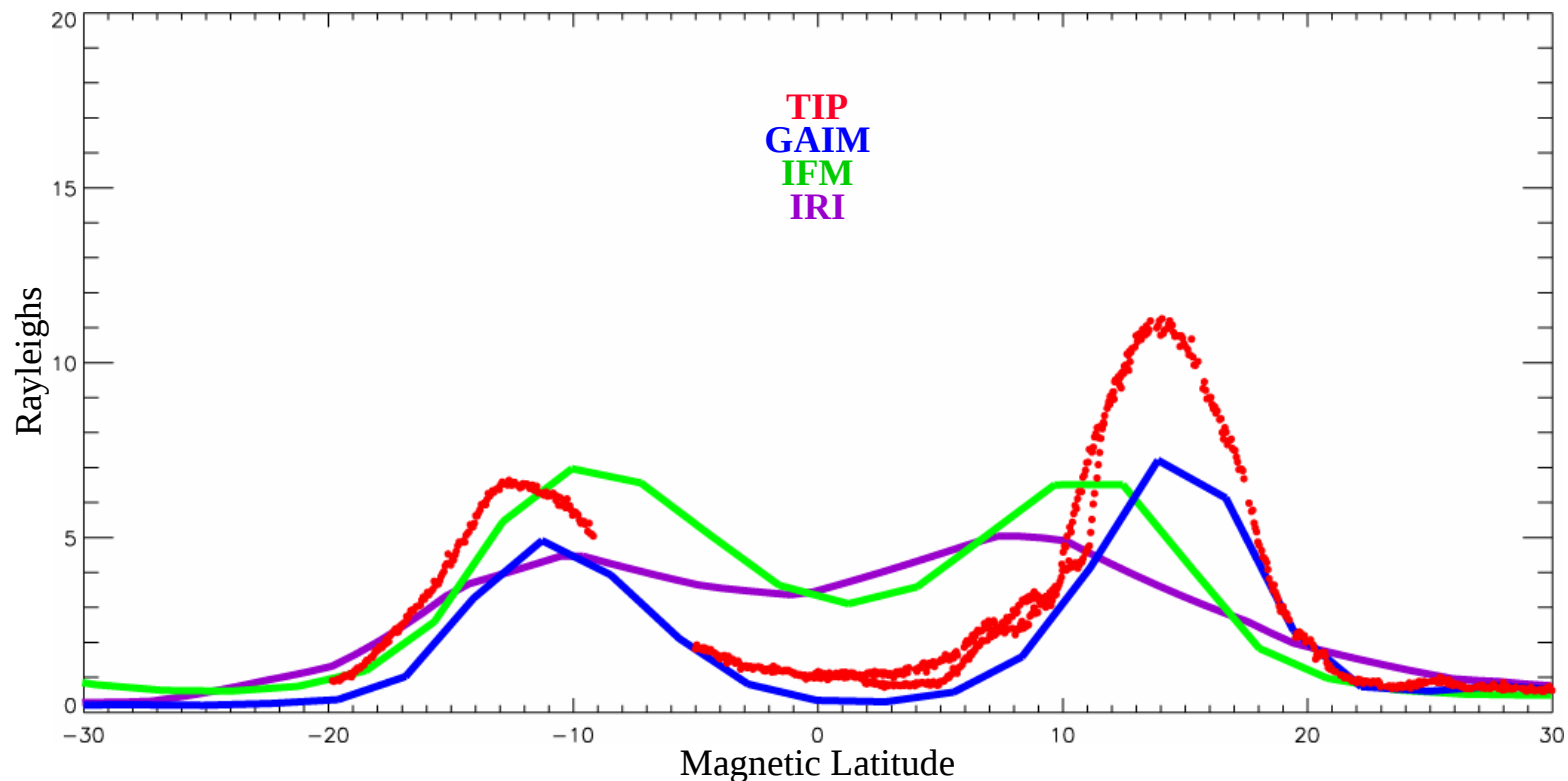




South American Pass



- TIP pass shows details of asymmetric anomaly
- GAIM captures location of crests and depth of trough, slightly underestimates crest magnitude
- Climatological models fail to accurately locate crests and depth of trough

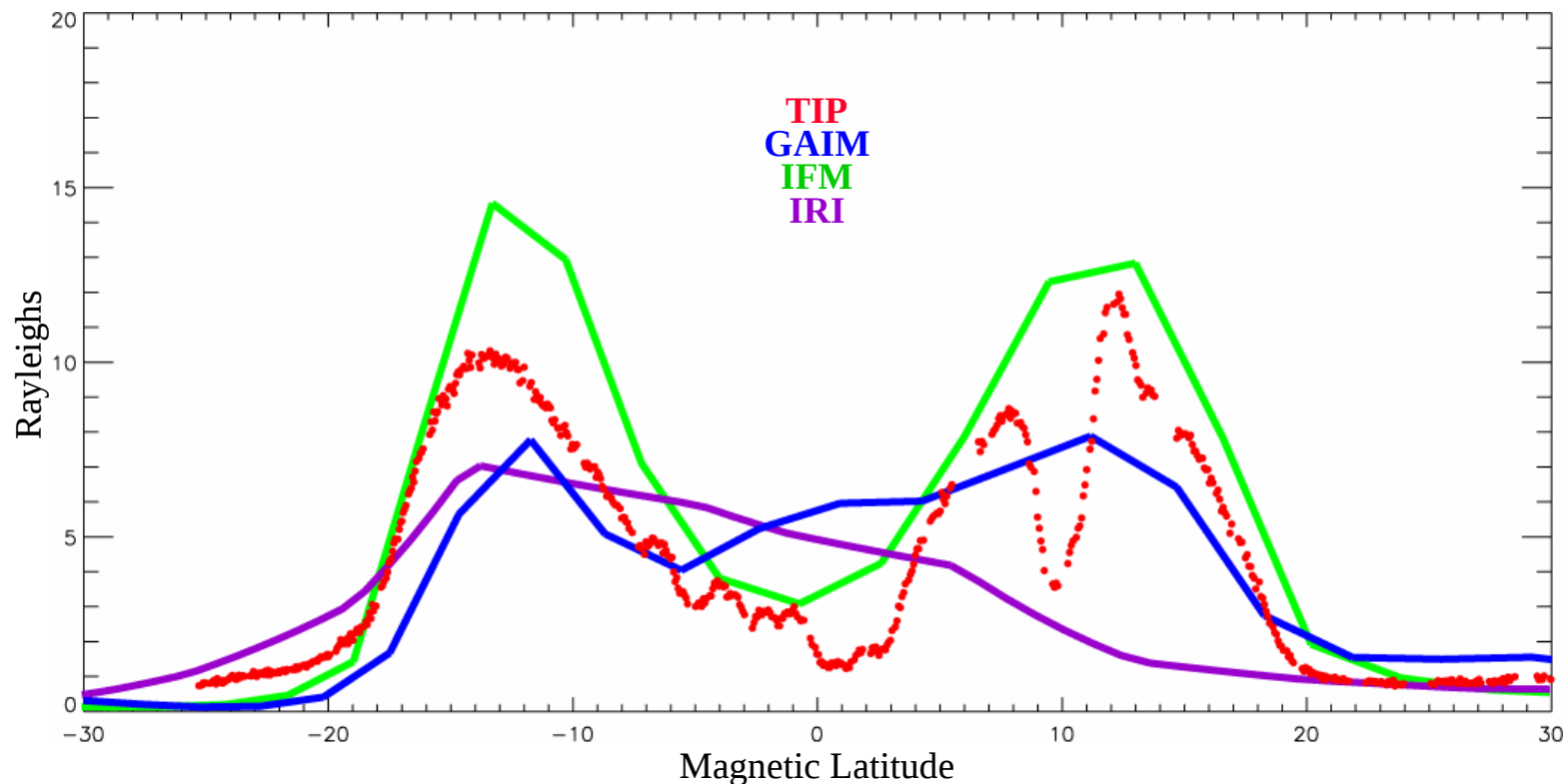




African Pass



- TIP pass shows anomaly modified by the presence of irregularity structures
- GAIM captures location of crests but not the depth of trough
- IFM accurately represents the crest location and depth of trough
- IRI doesn't show double crested anomaly or trough

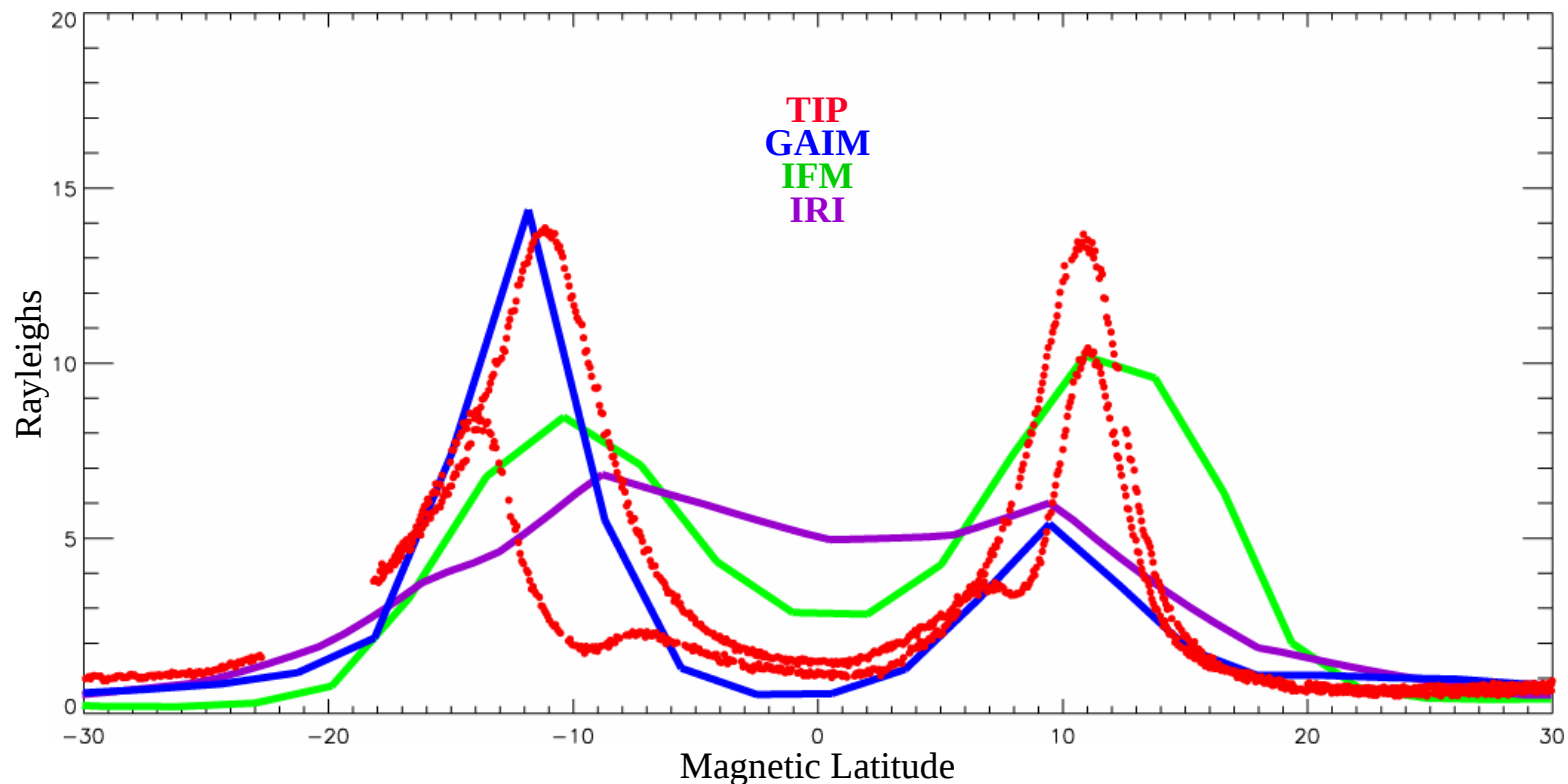




Hawaiian Sector Pass



- TIP pass shows symmetric anomaly and the presence of depletions
- GAIM captures location of crests and depth of the trough, underestimates northern crest
- IFM overestimates the width of the crests, underestimates depth of trough
- IRI shows only a hint of double-crested anomaly with almost no trough

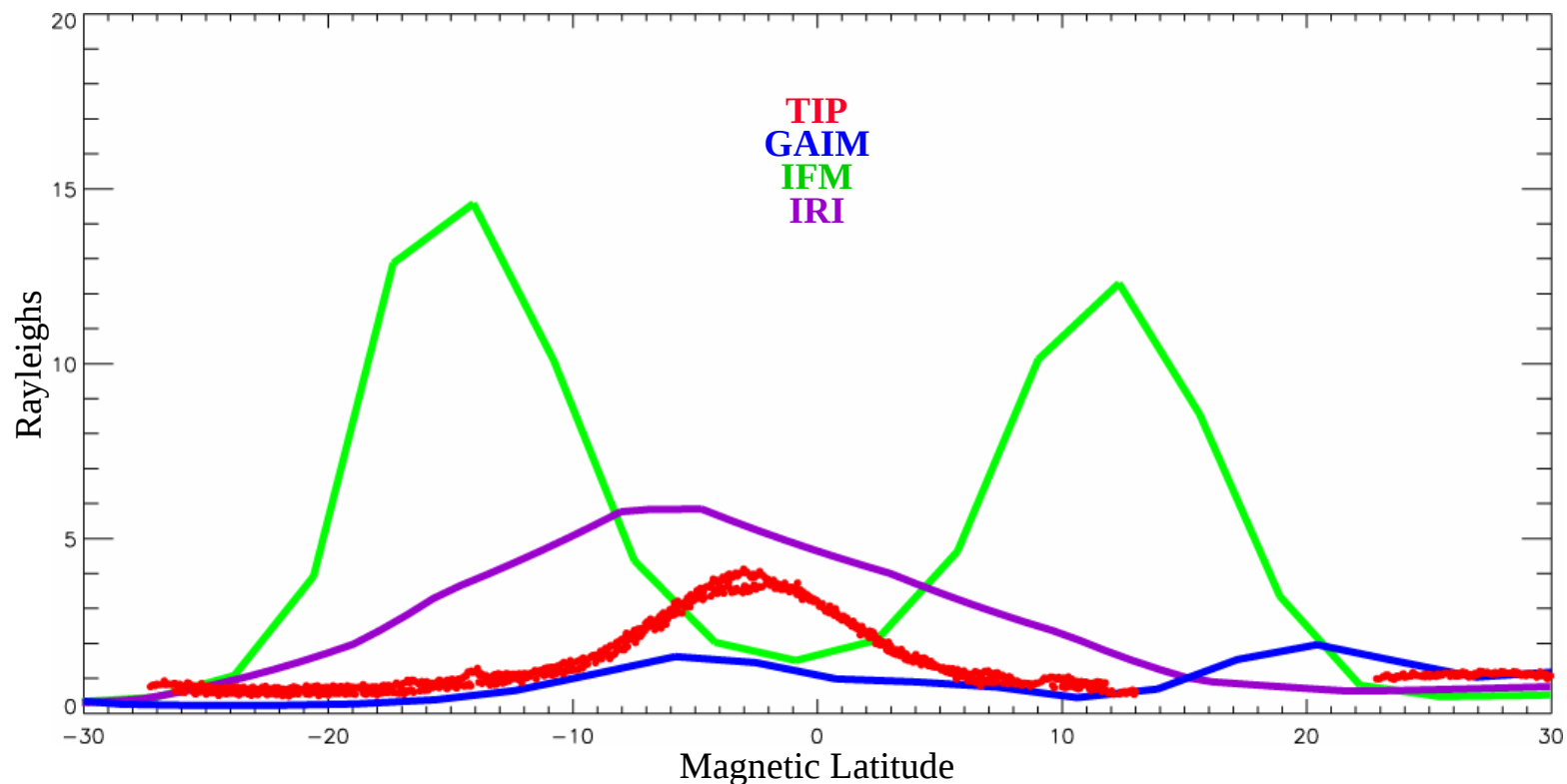




Indian Pass



- TIP pass shows absence of anomaly
- GAIM shows similar equatorial blob, although underestimated
- IFM does not capture this feature – still shows anomaly
- IRI's lack of anomaly is not a bad representation in this case, although much wider in latitude

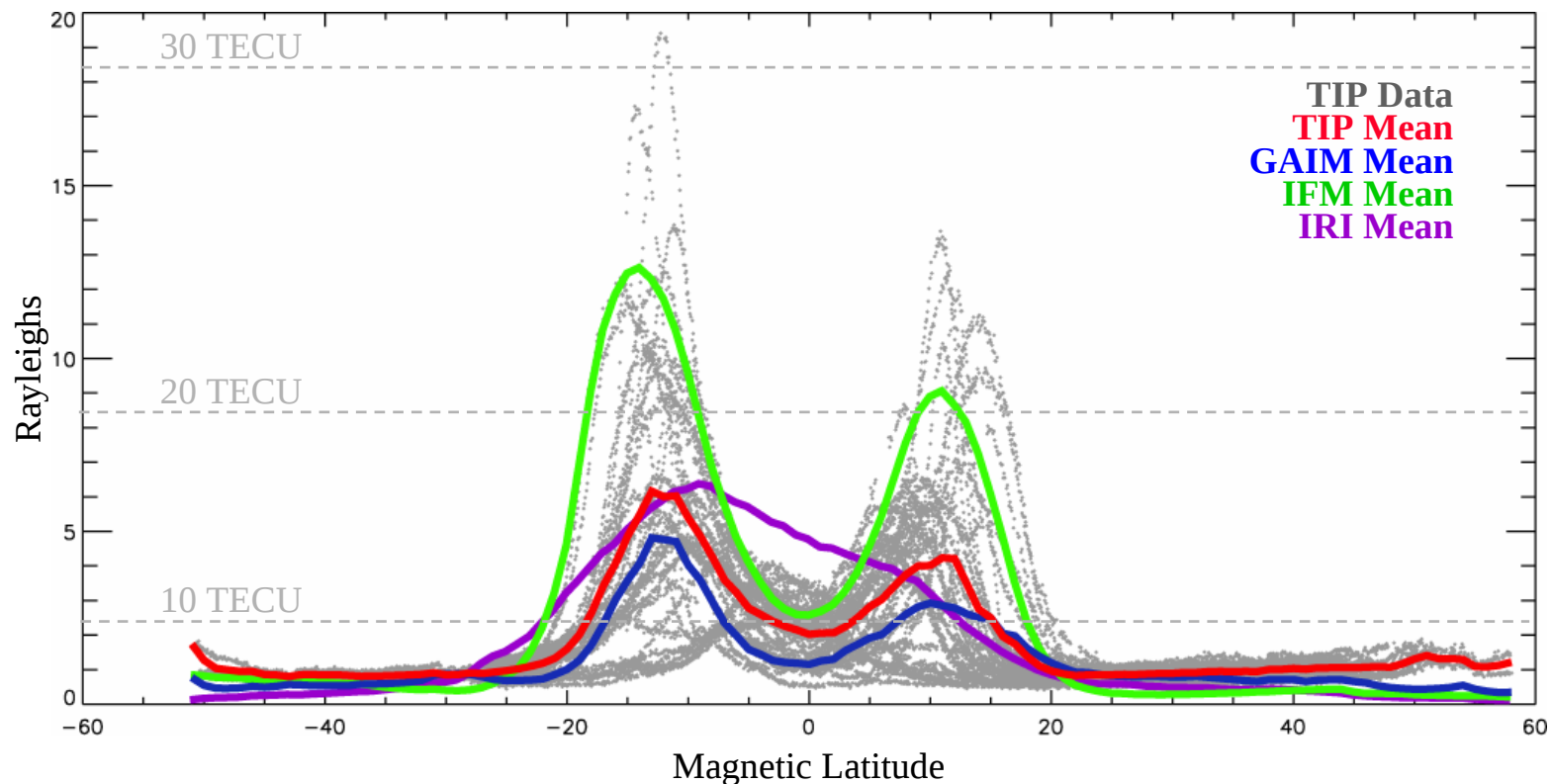




Average Equatorial Ionospheric Anomaly



- TIP data shows highly variable distribution of equatorial anomalies
- TIP average shows slight asymmetry in favor of Southern Hemisphere
- GAIM reproduces global anomaly with slightly less magnitude
- IFM routinely overestimates crest magnitude and width
- IRI doesn't show double crested anomaly or trough, reasonable mean

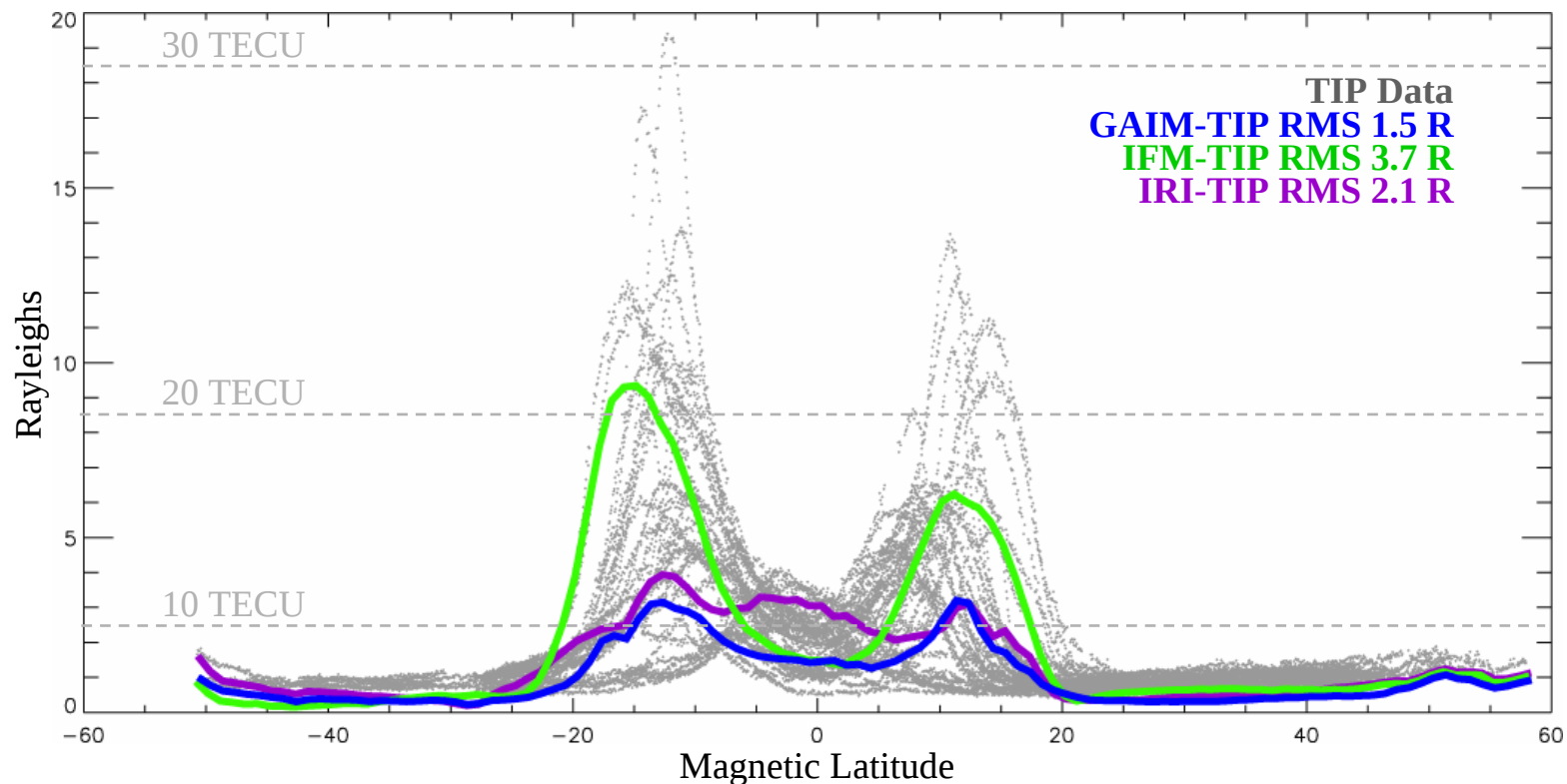




RMS Error as a Function of Magnetic Latitude



- GAIM shows significantly less error than IFM or IRI
 - Further error reduction, if source of bias identified
- IFM is statistically worse than IRI
- Even though IRI doesn't show an anomaly at 2100LT, on average it does better than IFM at the crests and worse than IFM in the trough





Conclusions



- TIP shows longitudinal variability of the post-sunset equatorial anomaly due to influence of non-migrating tides
- IFM and IRI models do not show similar tidal influence
- GAIMUSC captures the global variability of post-sunset equatorial anomaly
 - Assimilating ground-based GPS, and multiple RO sensors in same orbit (best case)
 - Morphology of equatorial anomaly is remarkably similar to TIP (qualitative)
 - Magnitude of the anomaly crests and trough are statistically better than IFM and IRI (quantitative)
 - Ingesting RO data helps specify the low latitudes more accurately in the horizontal direction!