

GLOBAL
LABORATORY
SYSTEMS



IONOSPHERIC EQUATORIAL ANOMALY STUDIES DURING SOLAR STORMS

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F. Walter

Brazil



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Summary

- Objective
- Magnetic storms
- VTEC evaluation
- Global VTEC maps developed
- Periods of magnetic storms analyzed and IEA behavior
- Conclusions and results

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Objective

- In order to study the global behavior of the **Ionospheric Equatorial Anomaly**, dynamic maps based on IONEX data have been generated.
- In this way the effects of the solar and geomagnetic activity on the Ionospheric Equatorial Anomaly can be easily correlated.

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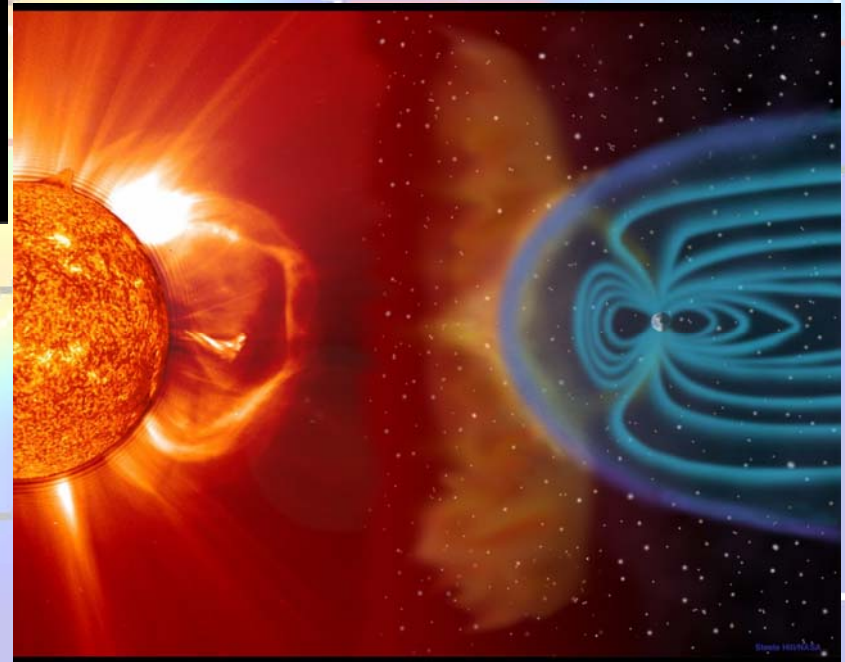
Coronal Mass Ejection (CME)



The Sun regularly sends massive solar explosions of radioactive plasma (coronal mass ejection - CME) with the intensity of a billion megaton bombs going through the solar system directly towards the Earth.



Credit: University of Colorado

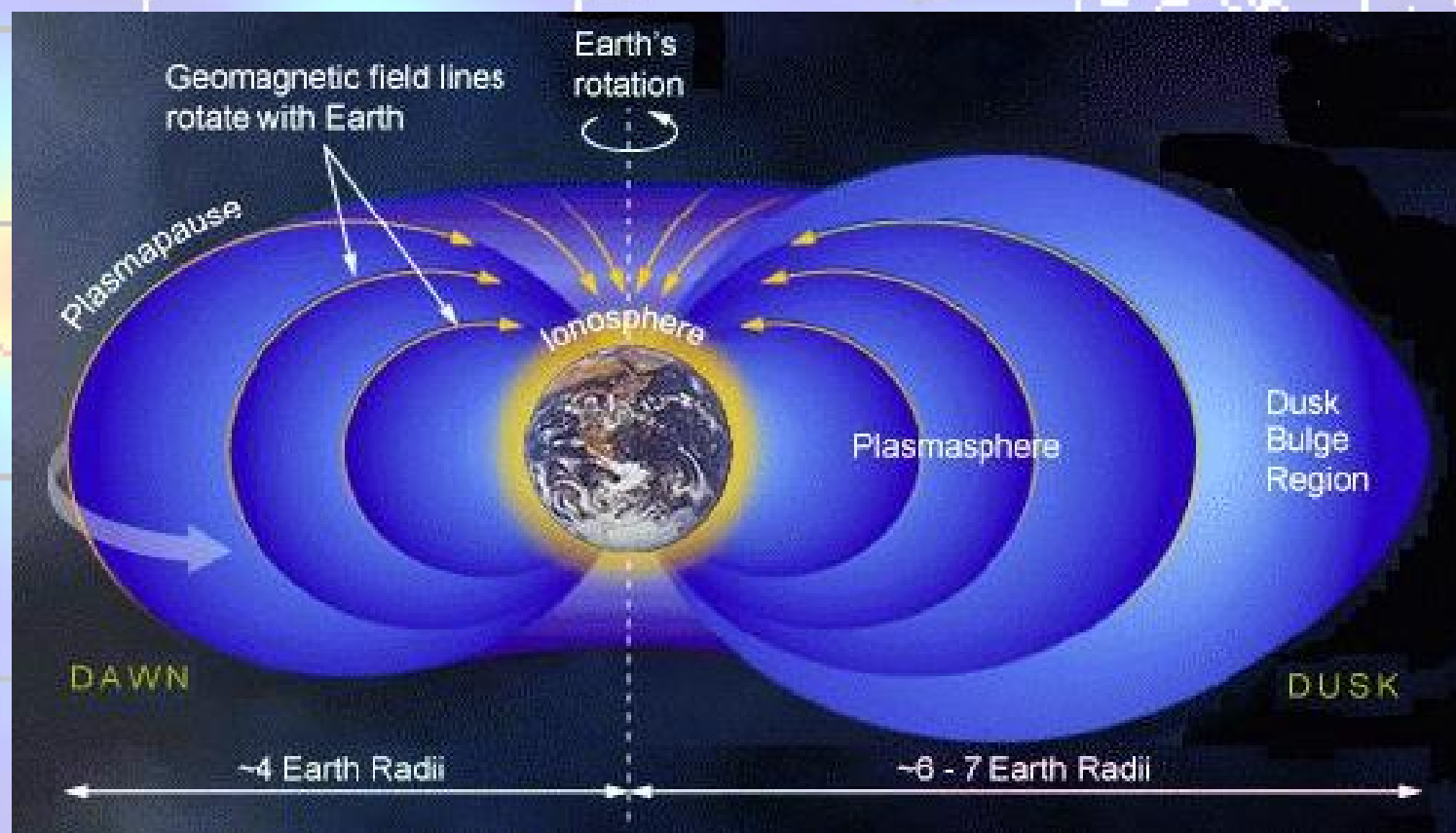


Credit: Nasa

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Ionosphere and Plasmasphere



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Magnetic Storms

Can cause effects like:

- bright auroras visible at more southern latitudes than usual; and
- even short out orbiting satellites, rendering them permanently inoperable.

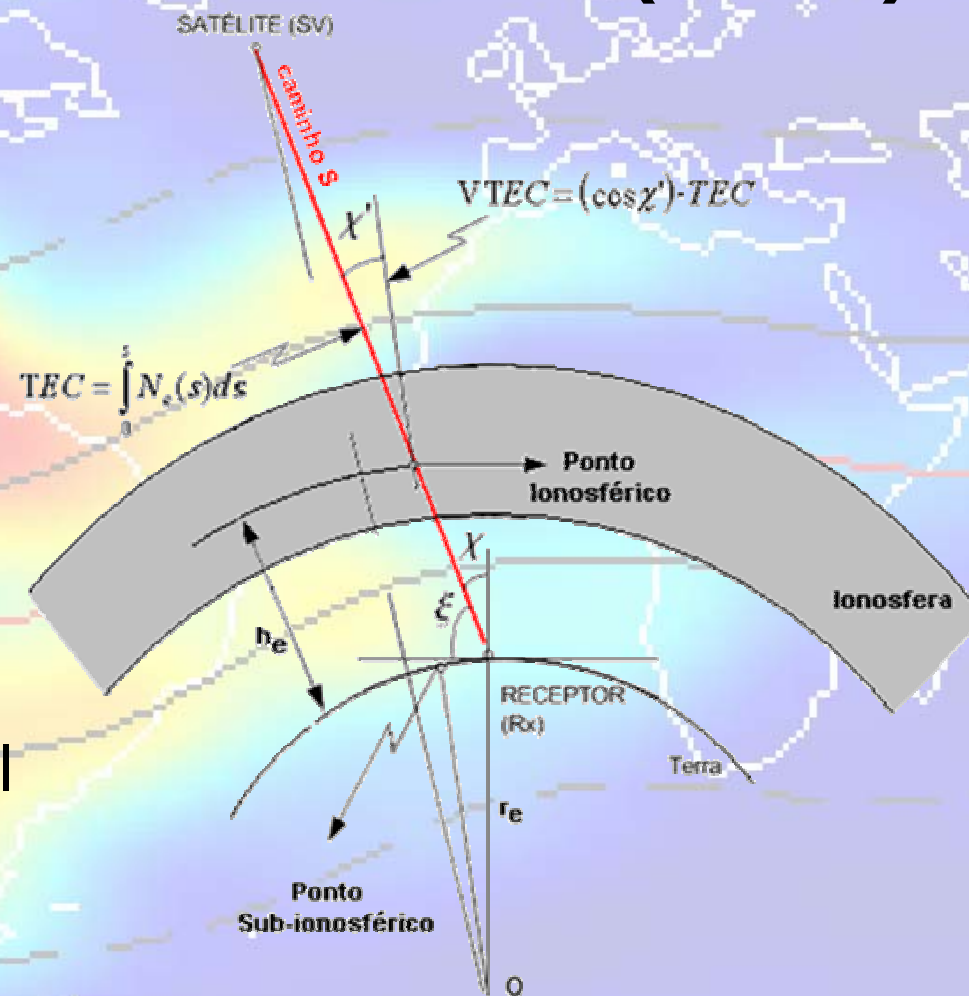
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Vertical Total Electron Content (VTEC)

Ionosphere is the major source of error in the user positioning determination by GNSS, mainly at low geomagnetic latitude regions ($\pm 20^\circ$) which are affected by the Ionospheric Equatorial Anomaly (IEA).



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Geomagnetic Equator

±40° geomagne

±20° geoe

Dark blue:

Dark red :

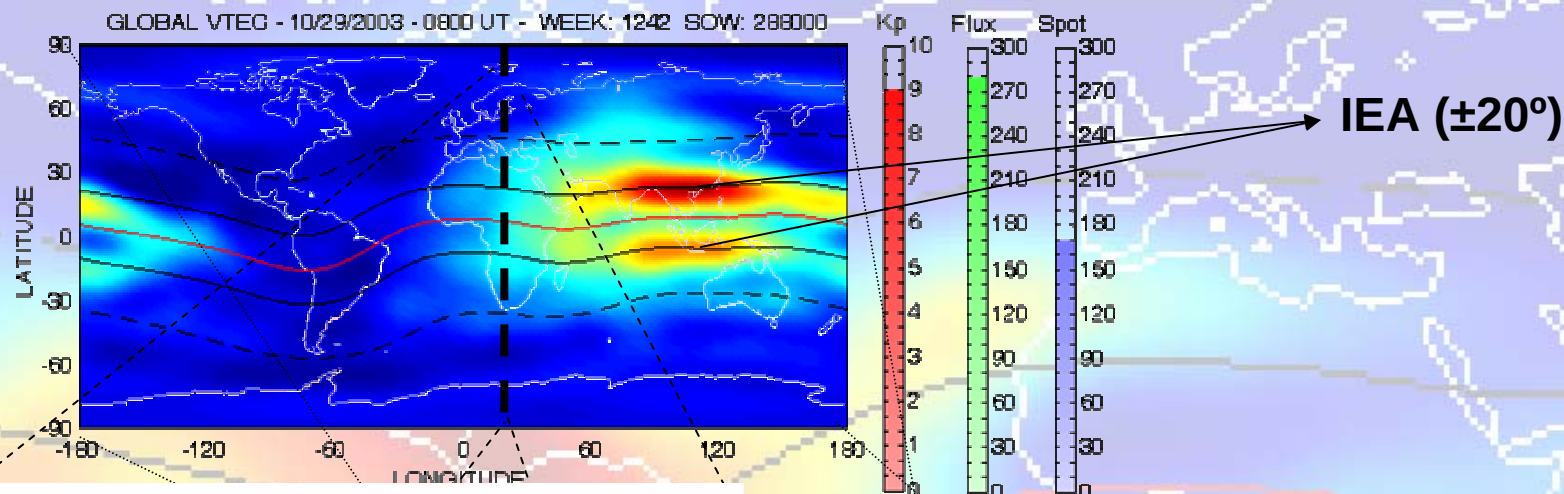
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Ionosp

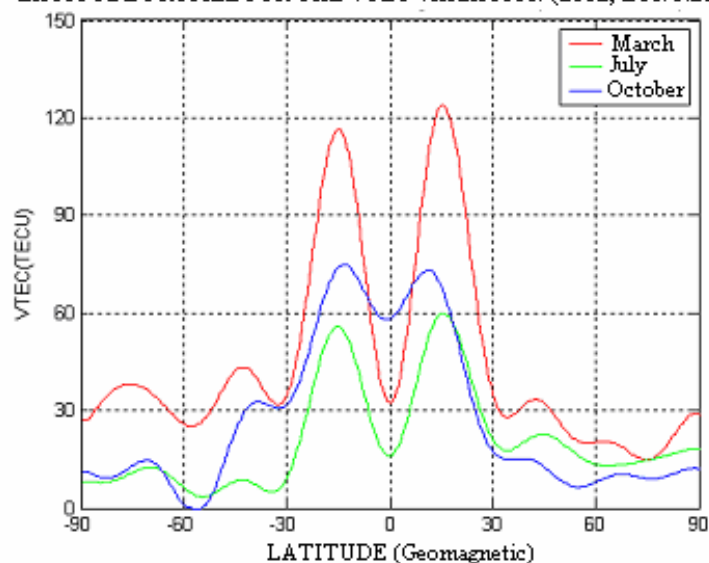




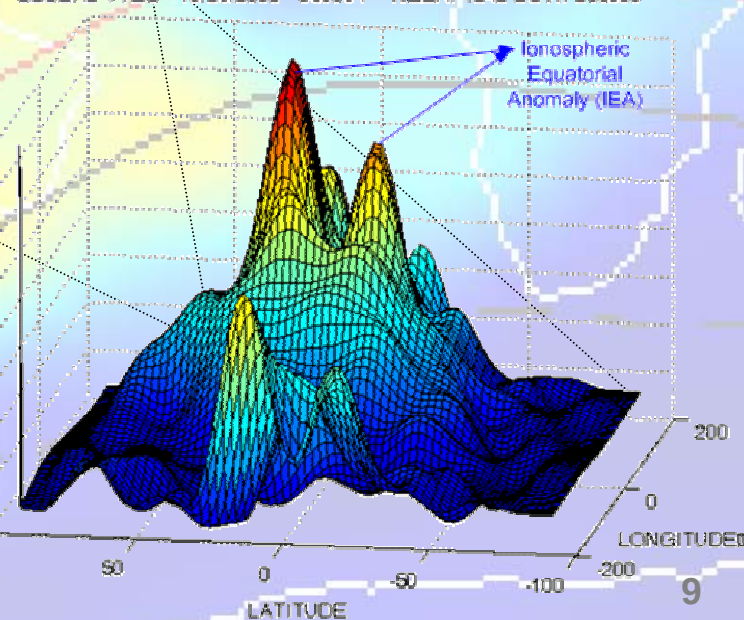
Ionospheric Equatorial Anomaly (IEA)



LATITUDE PROFILE FOR THE VTEC VARIATION (2002, LONG:200°)



GLOBAL VTEC - 10/29/2003 - 0800 UT - WEEK: 1242 SOW: 288000



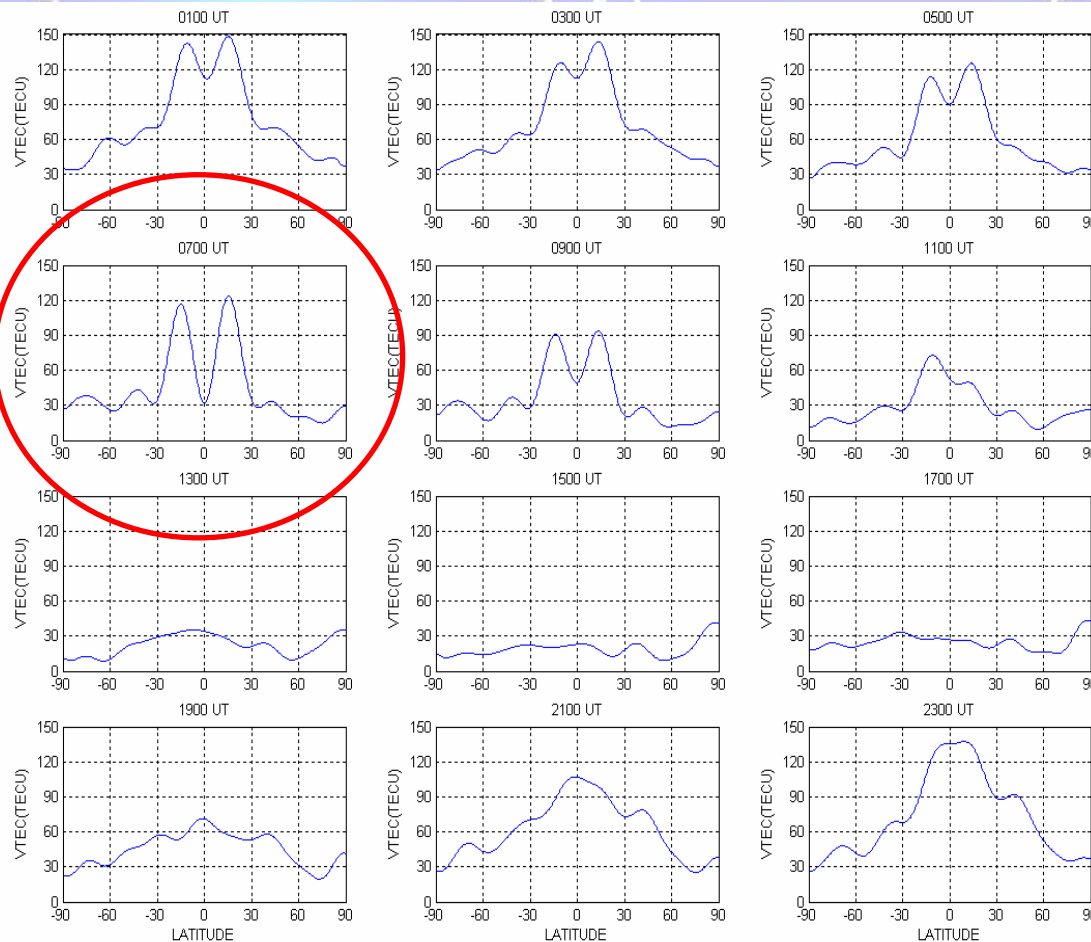
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IEA variation along the day (Long. 200°)



3 pm
Local Time



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Periods of magnetic storms



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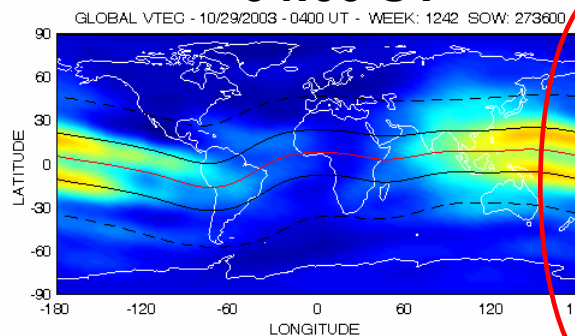
Periods of magnetic storms (October 29, 2003)



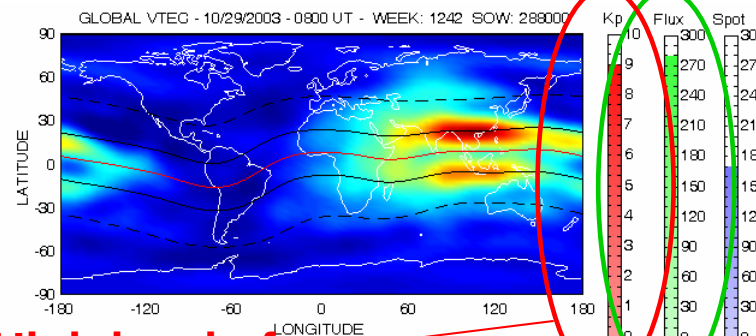
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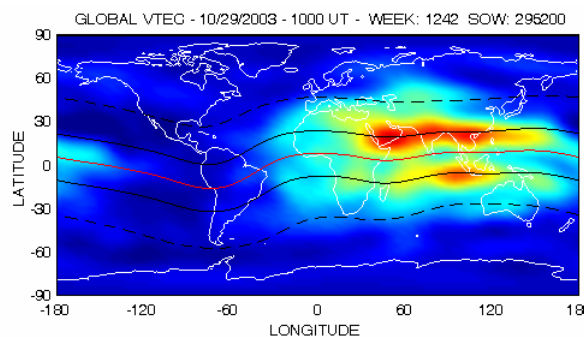
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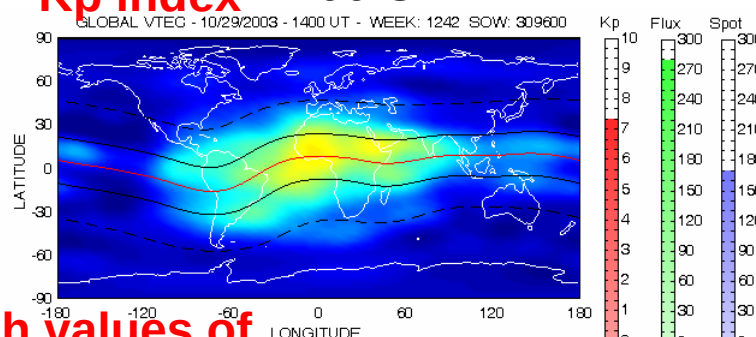
08:00 UT



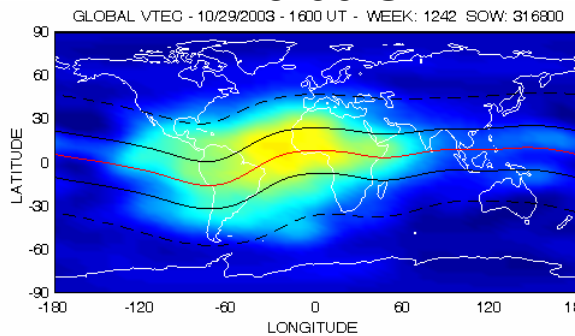
10:00 UT



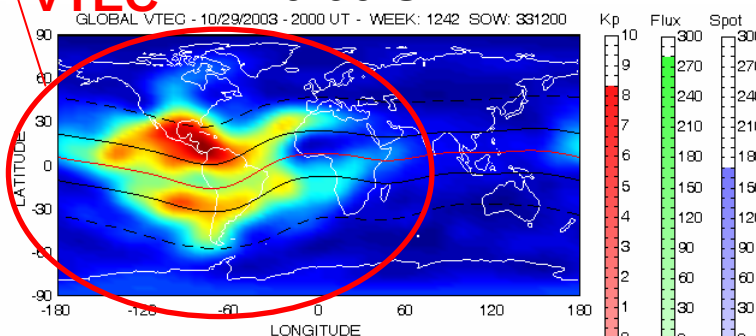
14:00 UT



16:00 UT



20:00 UT



0 20 40 60 80 100 120 140 160 TECU

0 20 40 60 80 100 120 140 160 TECU

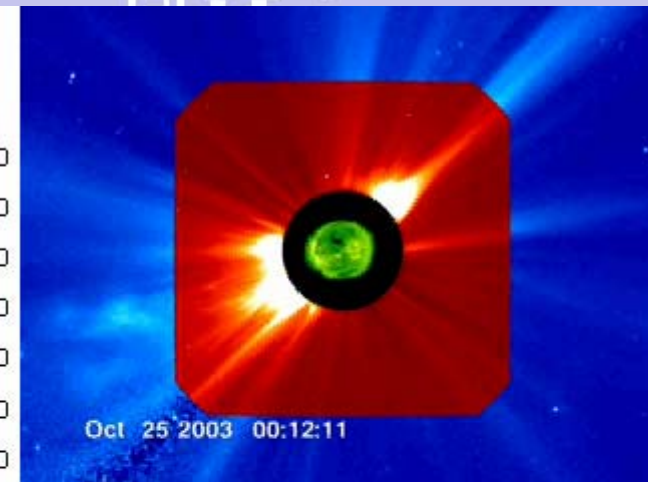
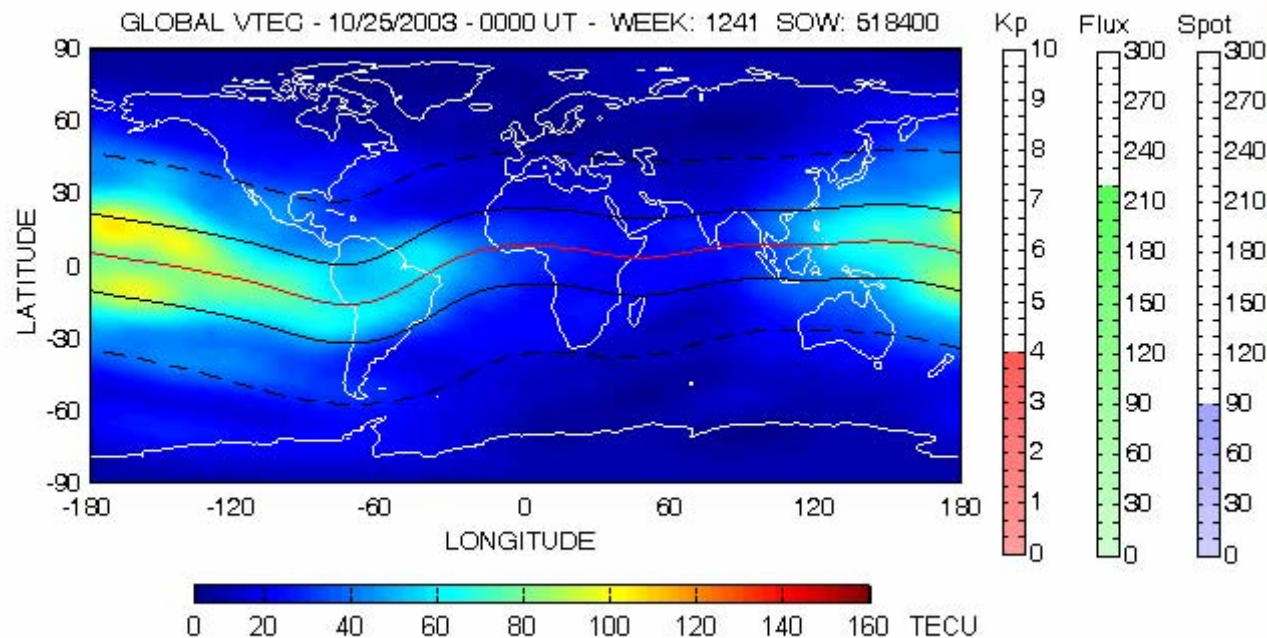
High level of
Kp index

High values of
VTEC



Periods of magnetic storms (October, 2003)

Credit: Nasa



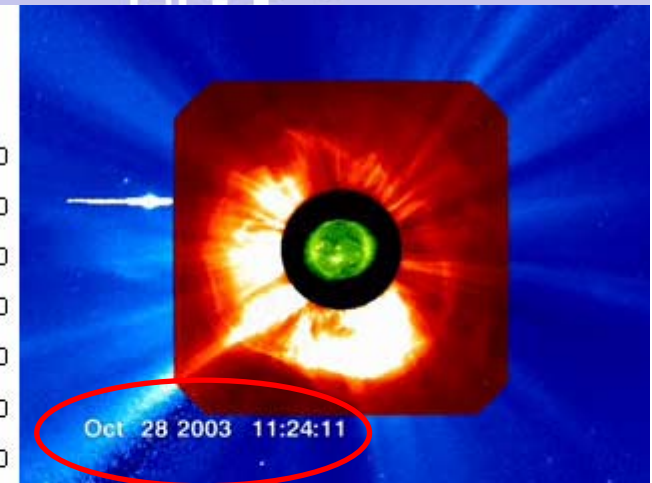
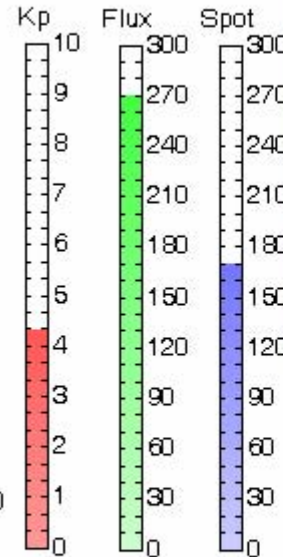
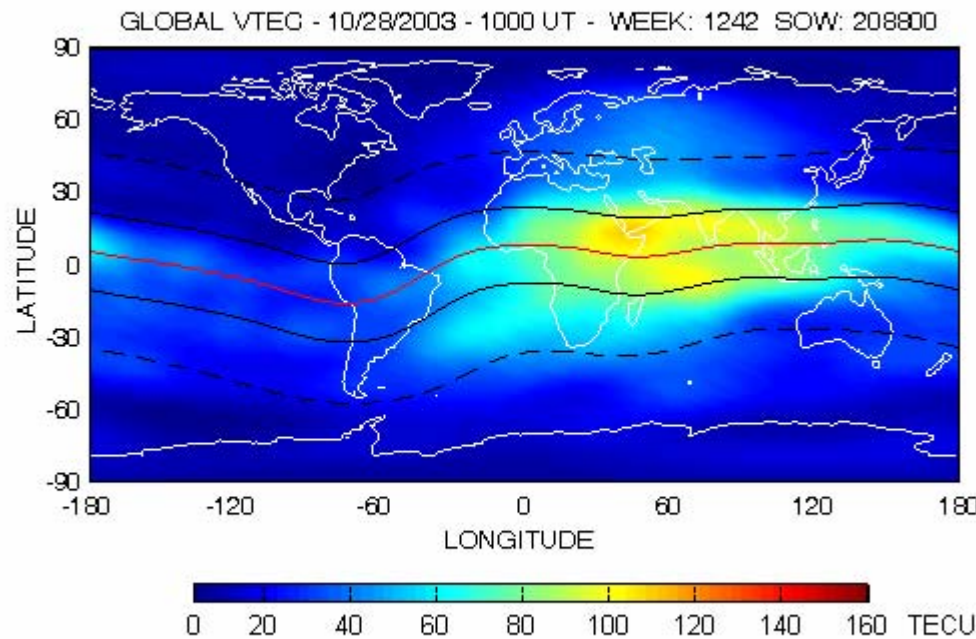
Copyright by





Periods of magnetic storms (October, 2003)

Credit: Nasa



**October 28, 2003
11:24 UT**

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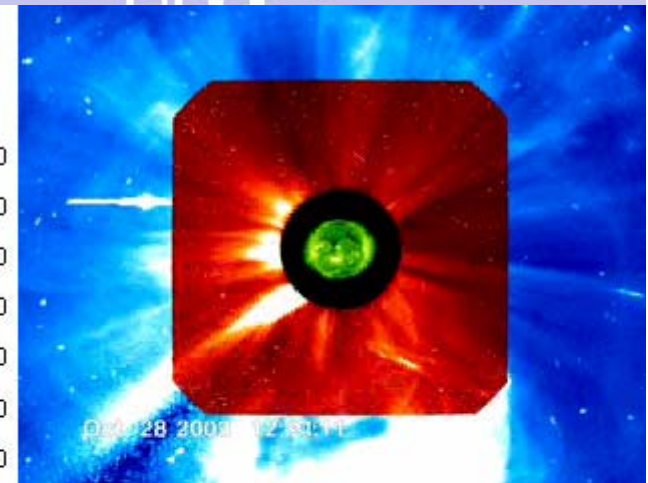
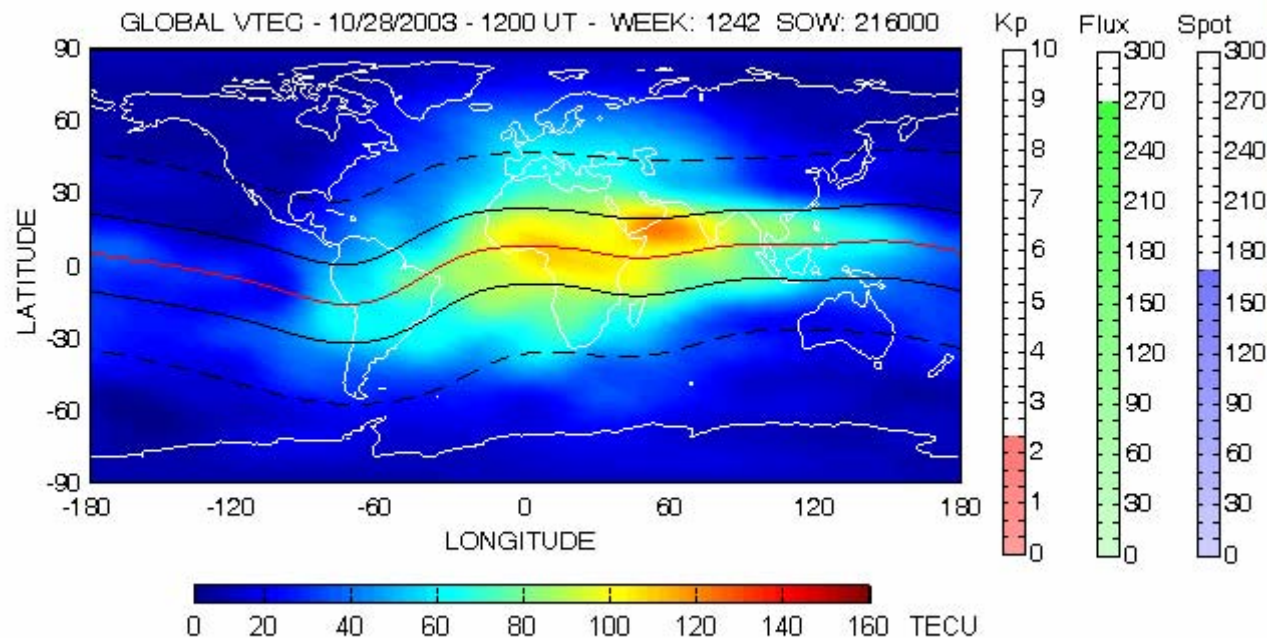
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Periods of magnetic storms (October, 2003)

Credit: Nasa



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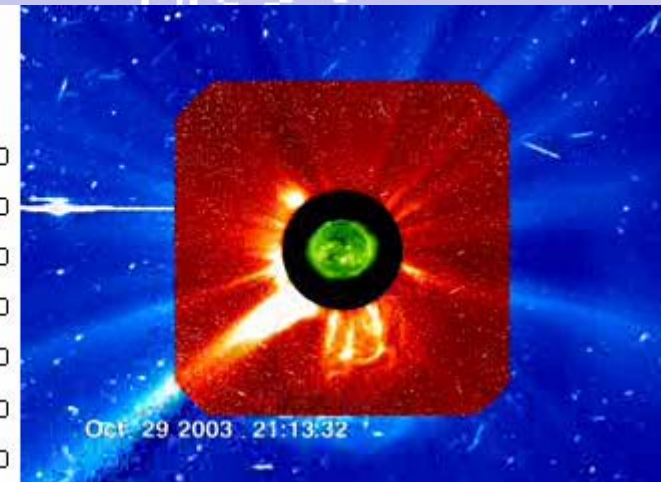
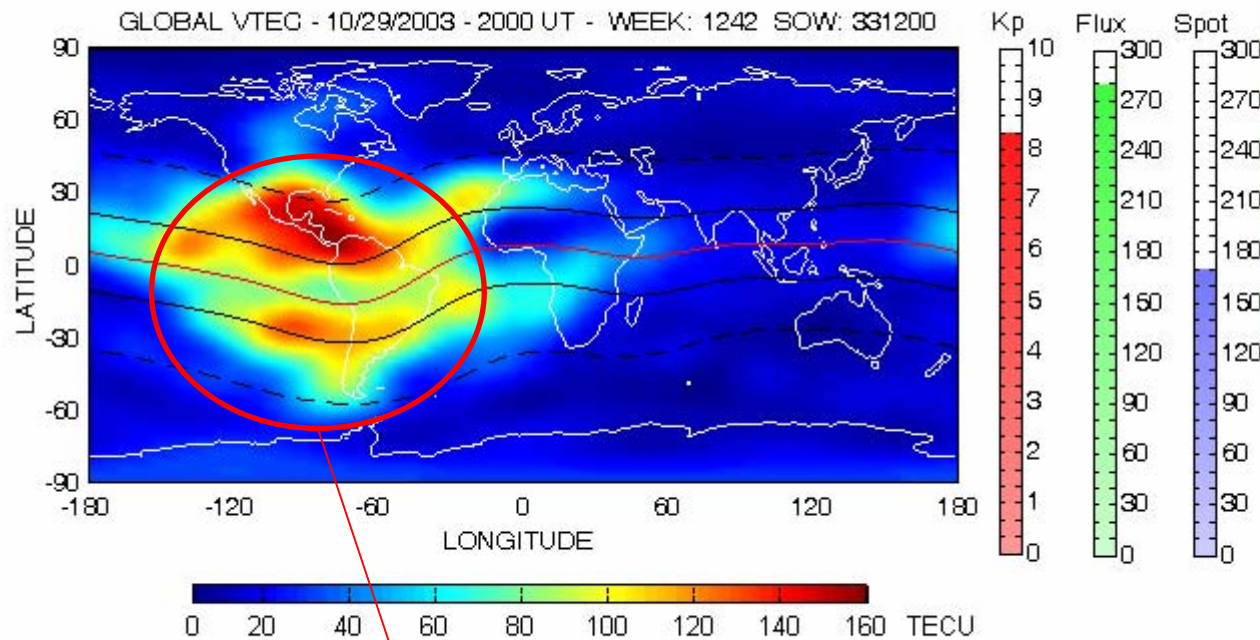
SKY
27/1/2006

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Periods of magnetic storms (October, 2003)

Credit: Nasa



October 29, 2003
20:00 UT

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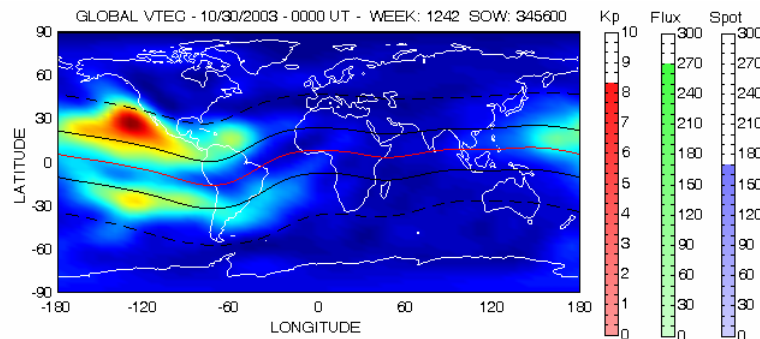
Periods of magnetic storms (October 30, 2003)



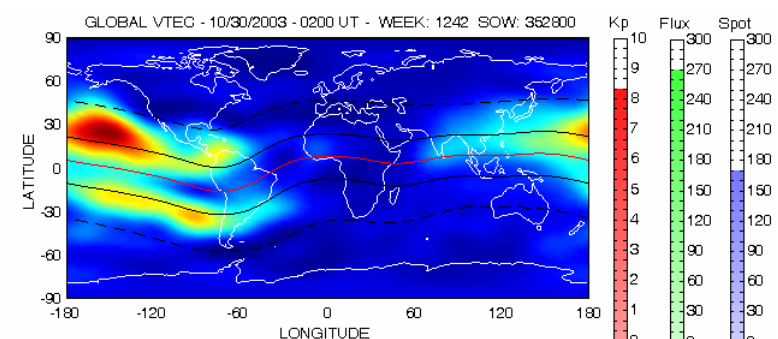
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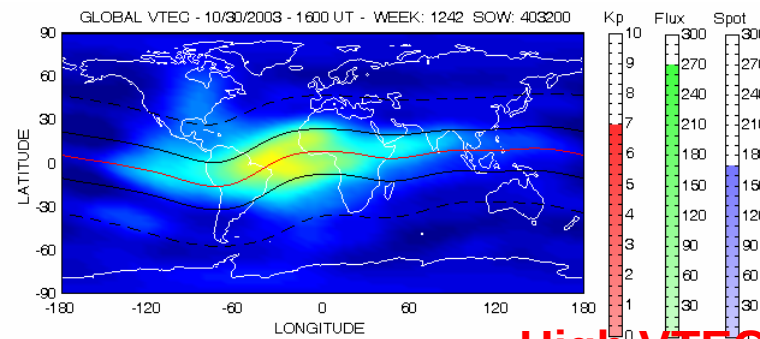
00:00 UT



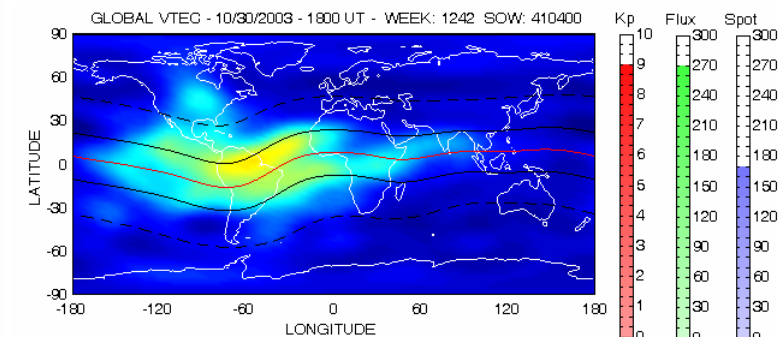
02:00 UT



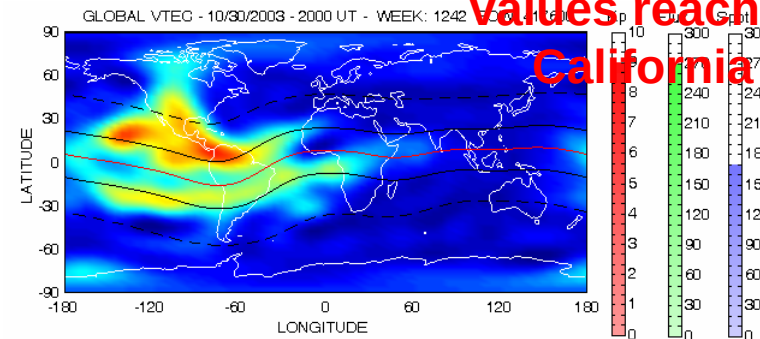
16:00 UT



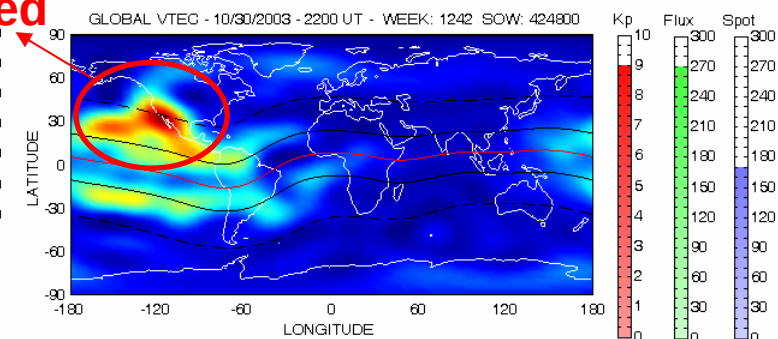
18:00 UT



20:00 UT



22:00 UT



High VTEC
values reached
California

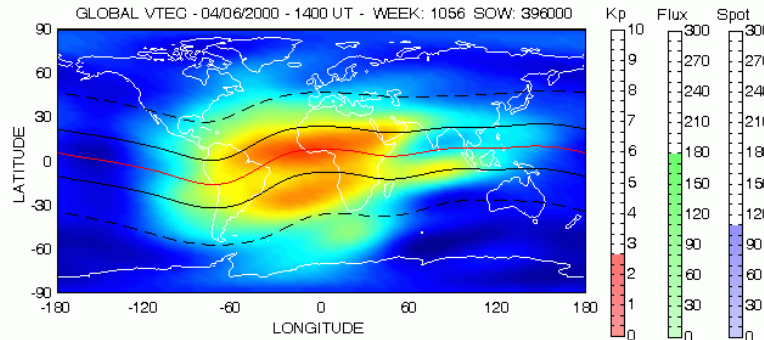
0 20 40 60 80 100 120 140 160 TECU

0 20 40 60 80 100 120 140 160 TECU

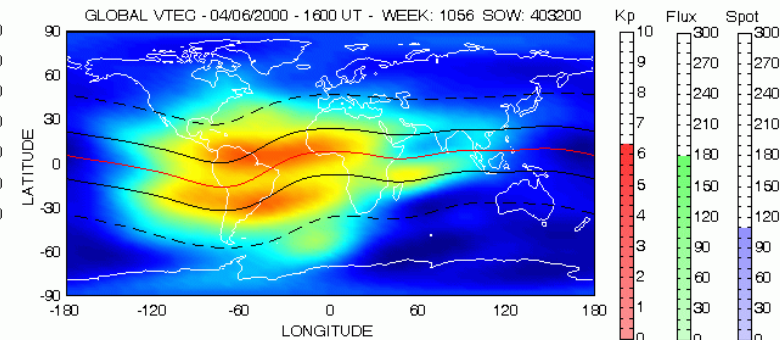
Periods of magnetic storms (April 06, 2000)



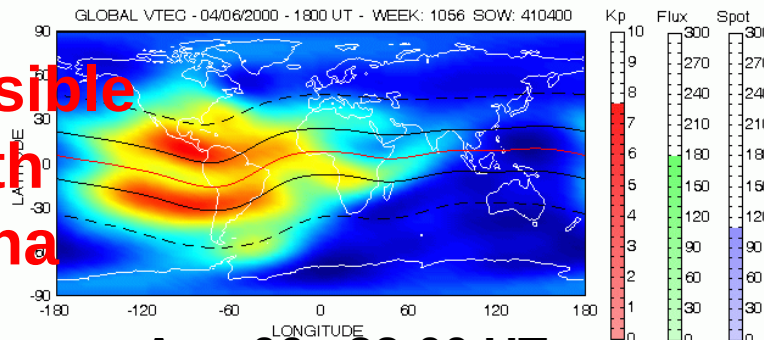
Apr. 06 - 14:00 UT



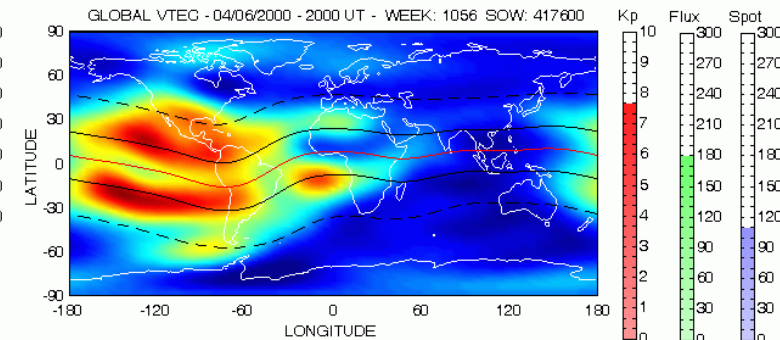
Apr. 06 - 20:00 UT



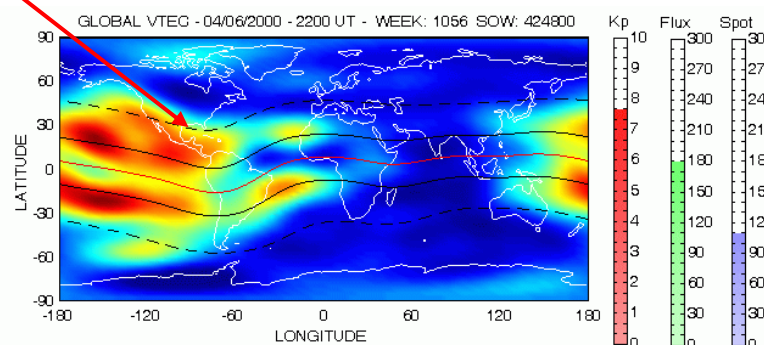
Apr. 06 - 18:00 UT



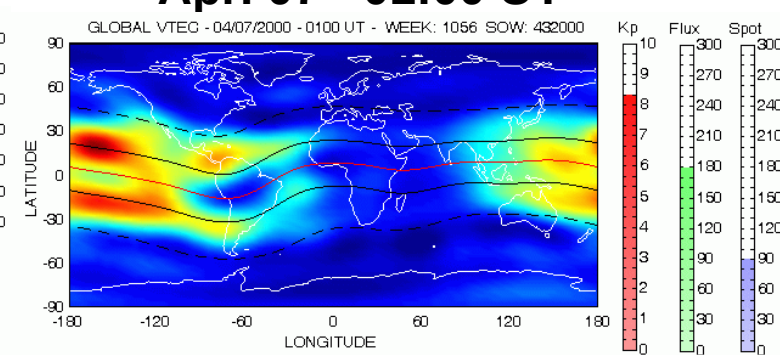
Apr. 06 - 20:00 UT



Apr. 06 - 22:00 UT



Apr. 07 - 01:00 UT



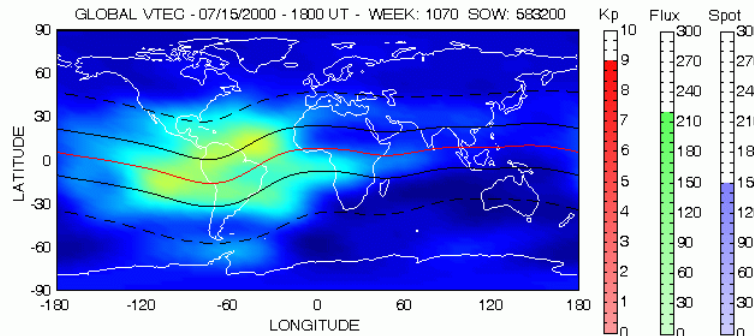
**Aurora visible
in North
Carolina**

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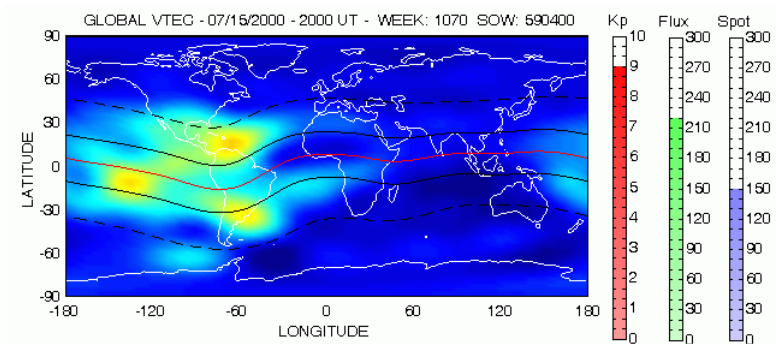
Periods of magnetic storms (July 15, 2000)



Jul. 15 - 18:00 UT

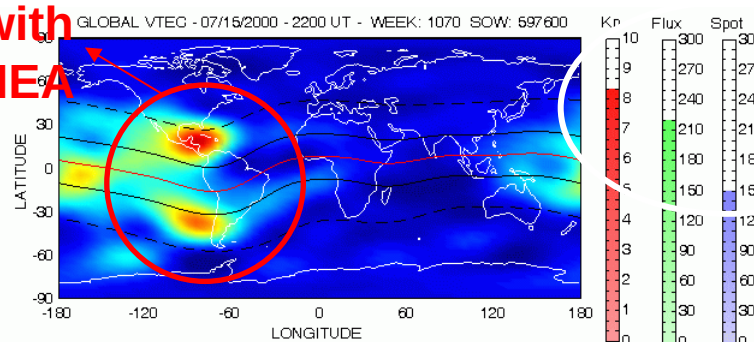


Jul. 15 - 20:00 UT

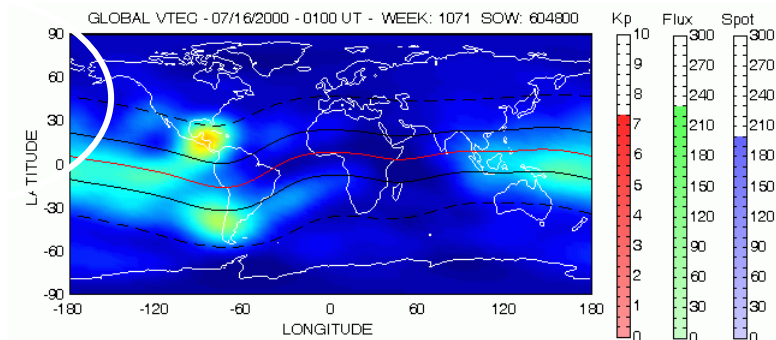


**High VTEC
values with
intense IEA**

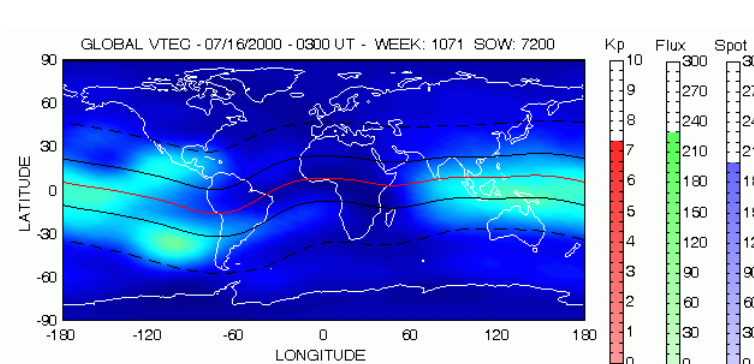
Jul. 15 - 22:00 UT



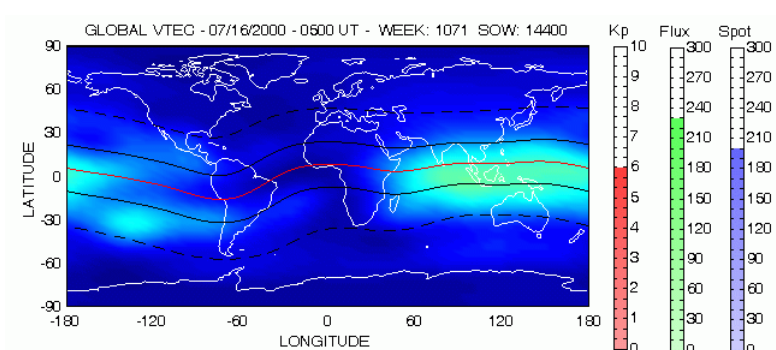
Jul. 16 - 01:00 UT



Jul. 16- 03:00 UT



Jul. 16- 05:00 UT



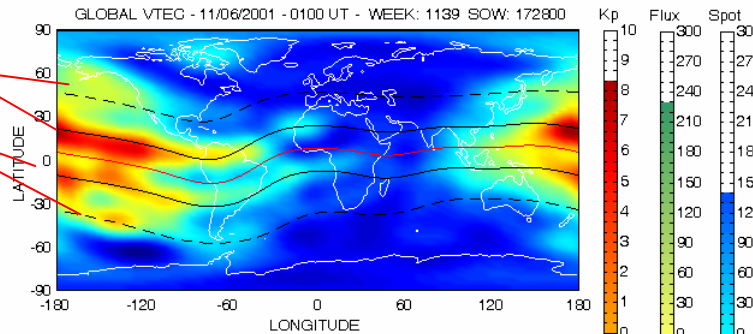
27/1/2006

Periods of magnetic storms (November 06, 2001)

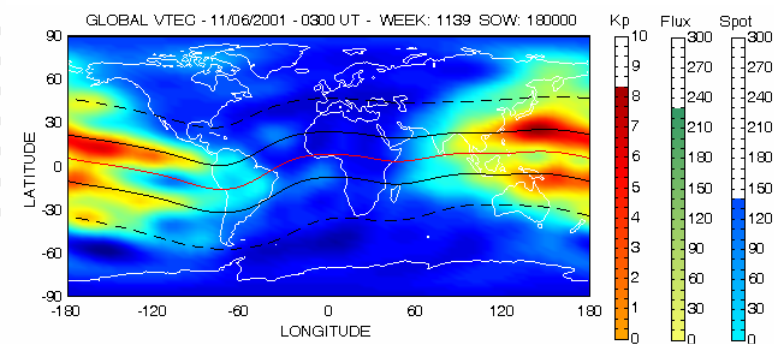


IEA intense
($\pm 20^\circ$ e $\pm 40^\circ$)

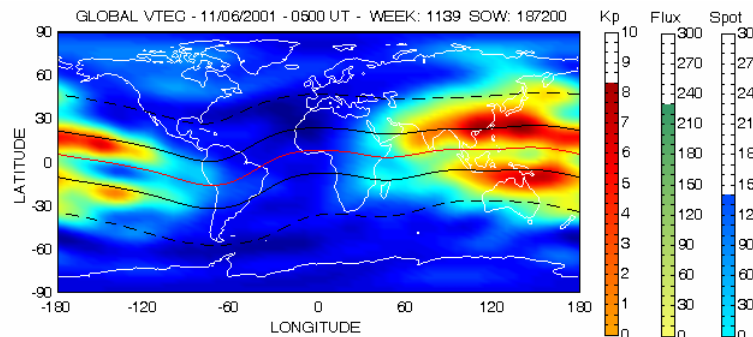
01:00 UT



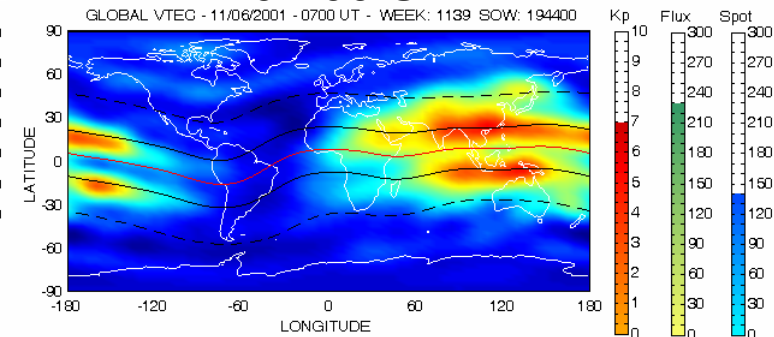
03:00 UT



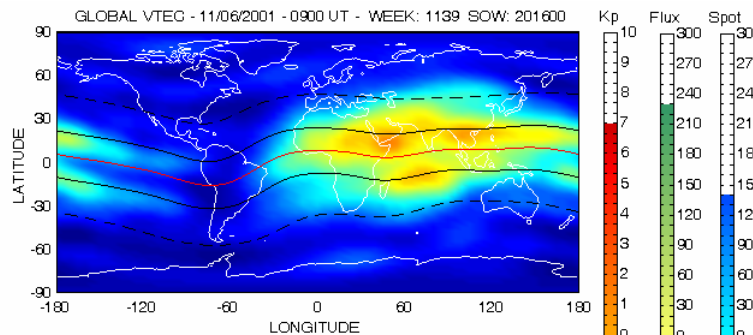
05:00 UT



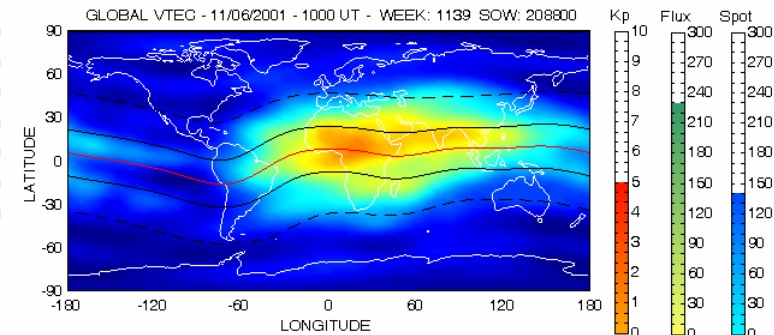
07:00 UT



09:00 UT



10:00 UT



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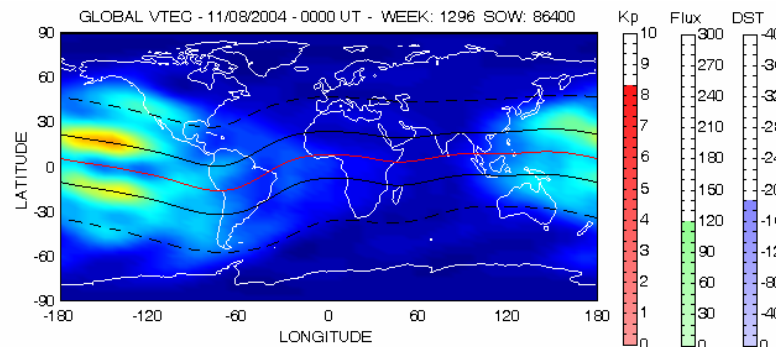
0 20 40 60 80 100 120 140 160 TECU

0 20 40 60 80 100 120 140 160 TECU

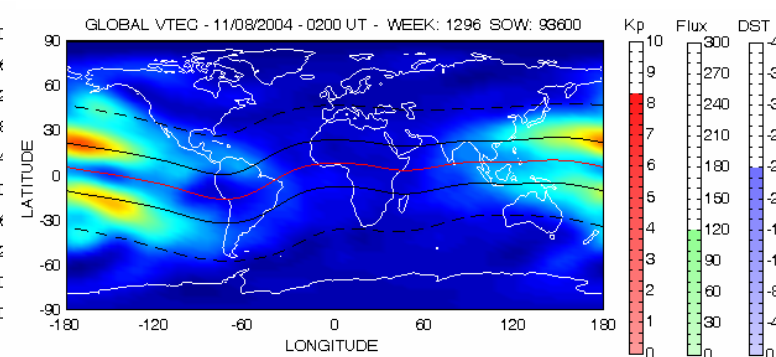
Periods of magnetic storms (November 08, 2004)



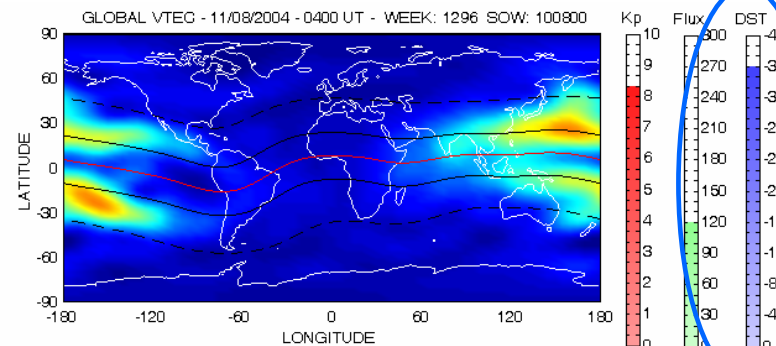
00:00 UT



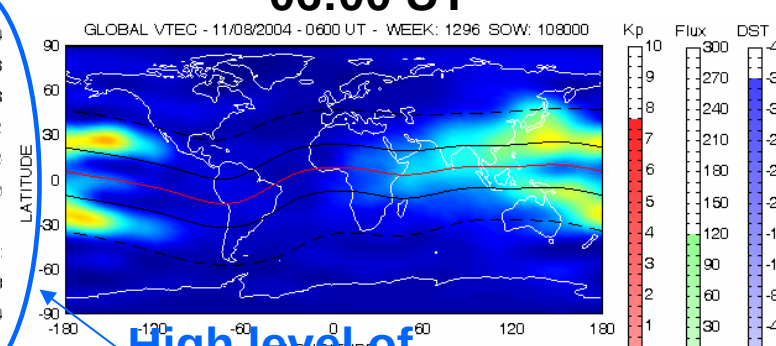
02:00 UT



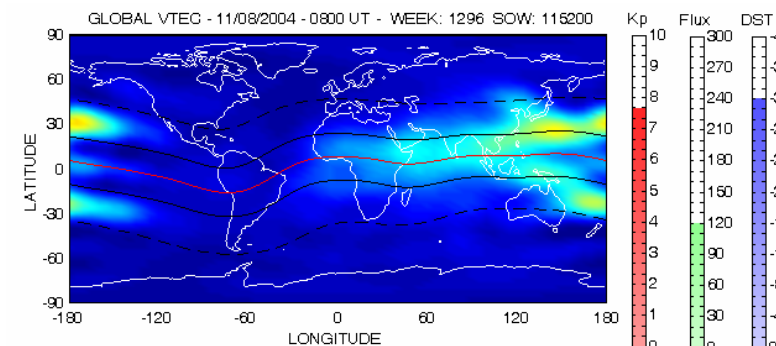
04:00 UT



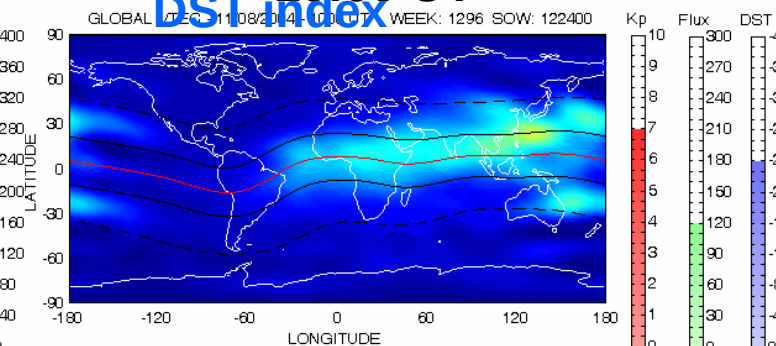
06:00 UT



08:00 UT



10:00 UT



High level of
DST index

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0 20 40 60 80 100 120 140 160 TECU

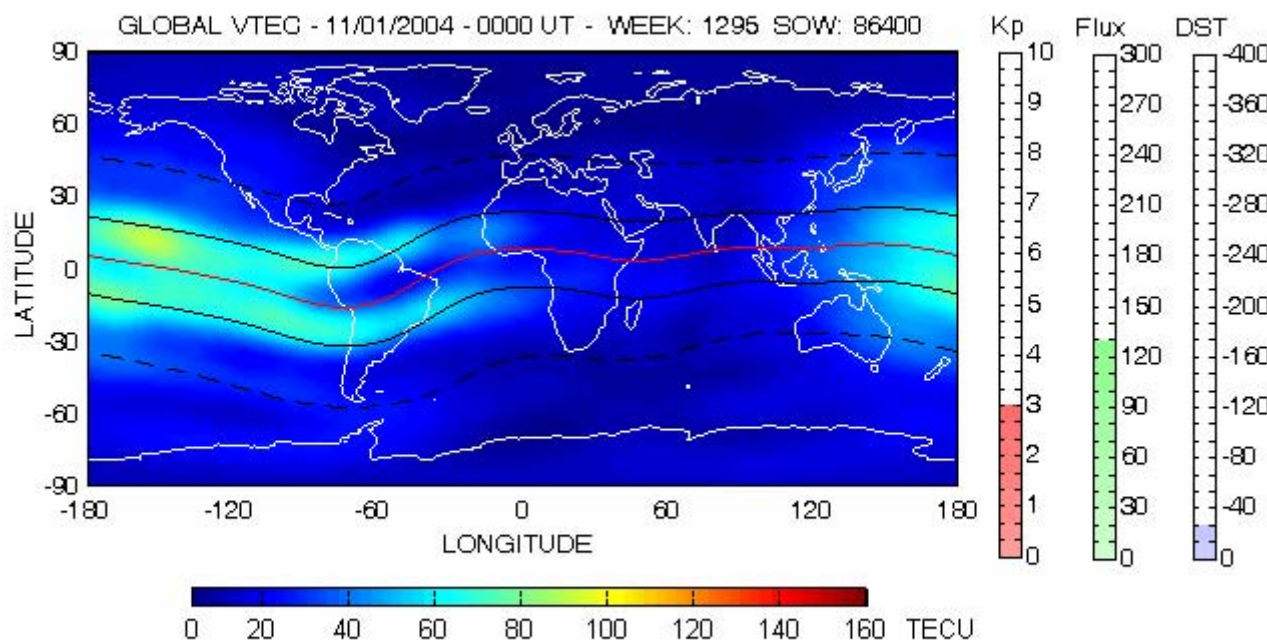
0 20 40 60 80 100 120 140 160 TECU

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Conclusions and Results

- This work has as a final product a database composed by VTEC maps since January 1999 which allows making comparisons with the geomagnetic activity, and the ionosphere behavior visualization during magnetic storms periods.



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Conclusions and Results

- This work has as a final product a database composed by VTEC maps since January 1999 which allows making comparisons with the geomagnetic activity, and the ionosphere behavior visualization during magnetic storms periods.
- Since we know, in at least two hours in advance, that what happens in region will affect another one, a global alarm system can be used.
- In order to study the GNSS error, it is necessary to take into account the ionospheric effects, mainly in the Equatorial Anomaly region or under high solar activity.
- The data analyzed show a delay of 20 to 30 hours (~1300 km/h) to the solar storm to affect the TEC. More data needs to be analyzed.
- A new data are starting to be used: the Dst.

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FURTHER RESULTS

- Since some discussion between IRI model and TEC were under taking during this meeting we show some preliminary results were the IRI gave us a guidance to model the global VTEC.
- (The following results are an extension of this work, it is only an advance of results getting by Thiago Moraes, a student of the GNSS Laboratory, as part of his Master Theses)

PERFIL DA DENSIDADE ELETRONICA EM FUNÇÃO DA LATITUDE

Epoca

Ano: 2001

Mes: Out

Dia: 30

Hora: 17 UT

Plotar dados IONEX

Plotar Perfil Dens Eletronica

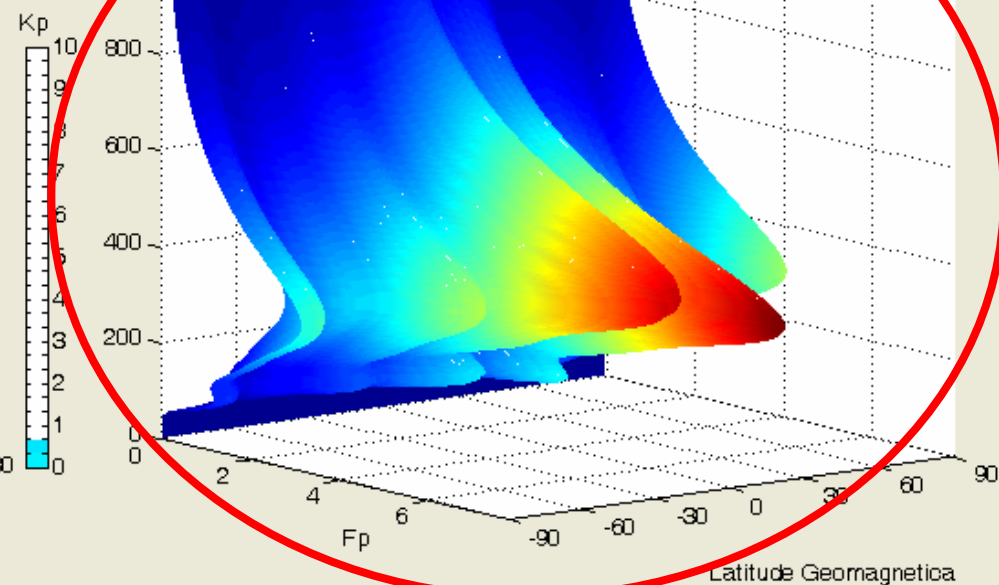
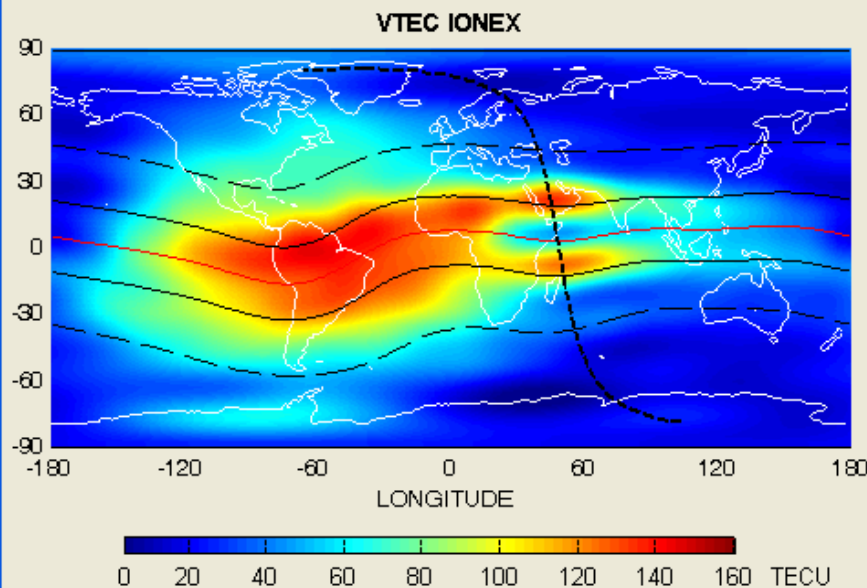
Sair

Coordenadas geograficas

Latitude: 6.61

Longitude: 48.19

Atualizar interface



COMPARAÇÃO VTEC IONEX x VTEC TRAÇADO DE RAIOS

Epoca

Ano: 2001

Mes: Out

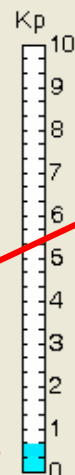
Dia: 30

Hora: 17 UT

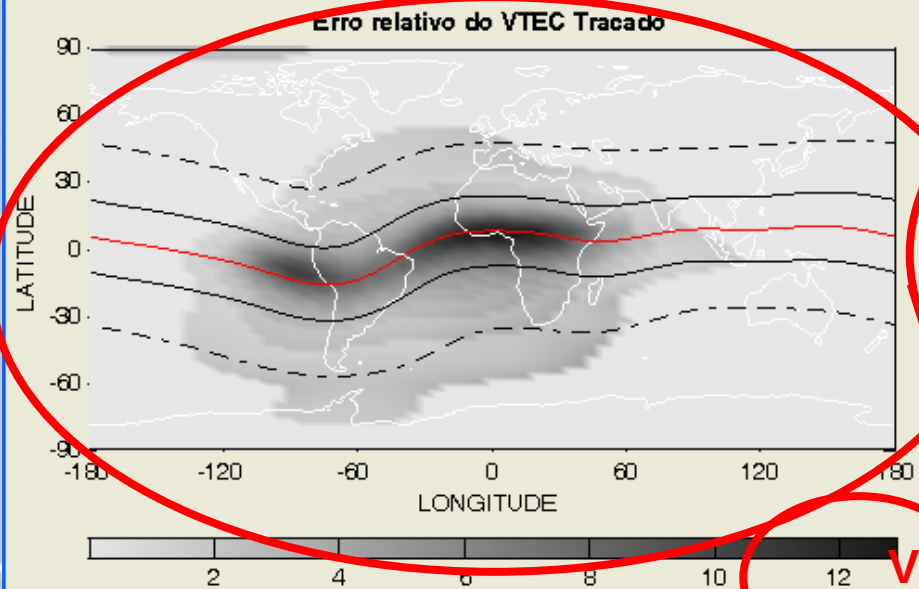
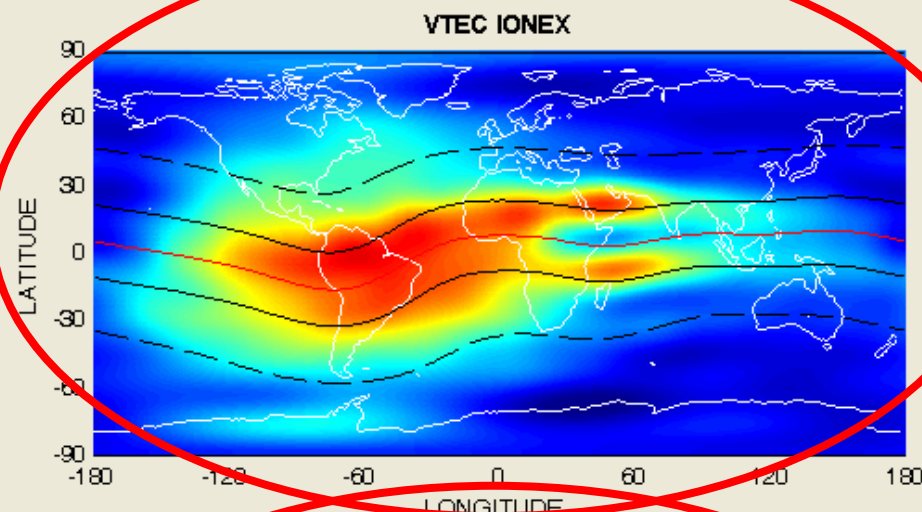
Plotar dados IONEX

Plotar VTEC do Traçado de raios

Plotar Diferença

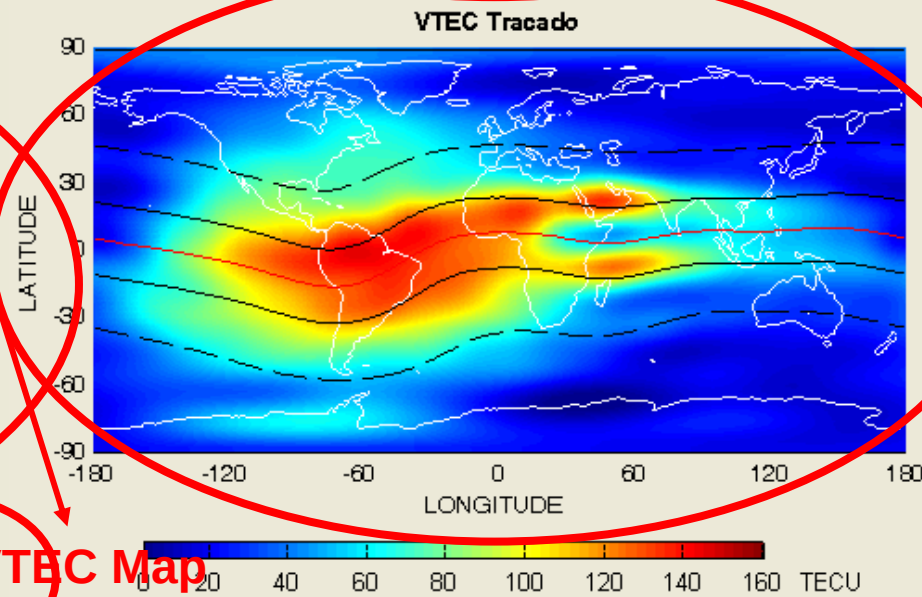


VTEC Map
using density profile
based on IRI parameters



Relative error:
 $< 15 \times 10^{-3} \%$

VTEC Map
using density profile
based on IRI parameters



PERFIL DA DENSIDADE ELETRONICA EM FUNÇÃO DA LATITUDE

Epoca

Ano: 2003

Mes: Out

Dia: 30

Hora: 22 UT

Plotar dados IONEX

Plotar Perfil Dens Eletronica

Sair

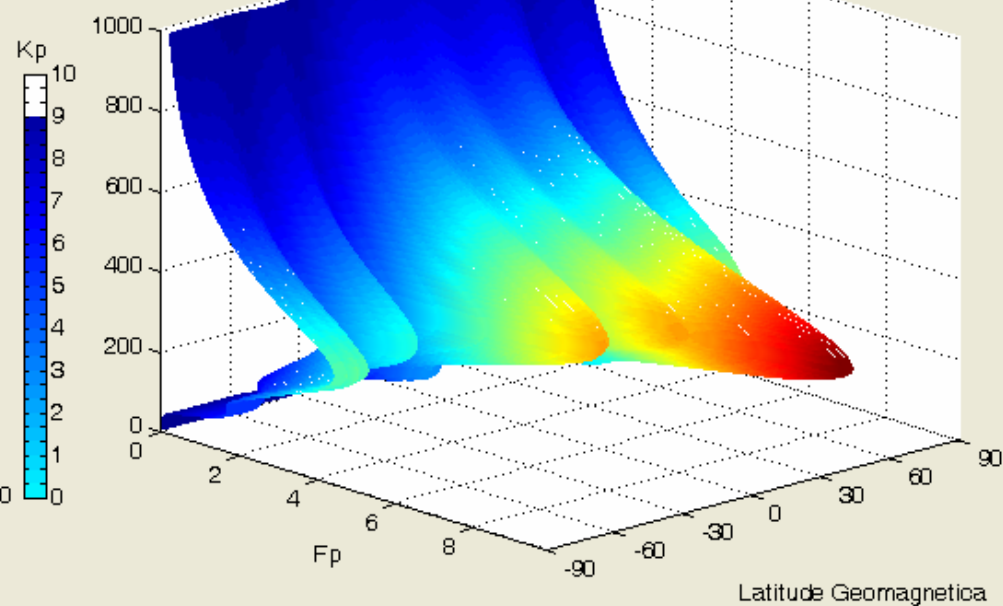
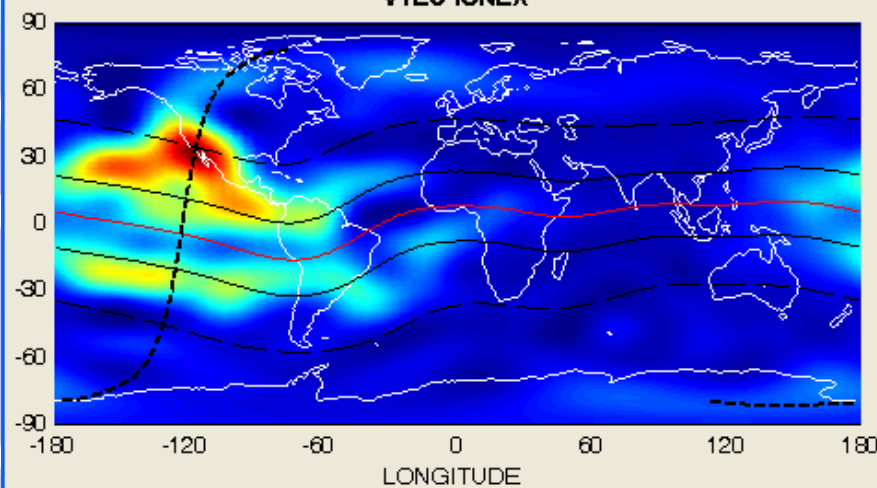
Coordenadas geograficas

Latitude: 29.72

Longitude: -117.39

Atualizar interface

VTEC IONEX



COMPARAÇÃO VTEC IONEX x VTEC TRAÇADO DE RAIOS

Epoca

Ano

Mes

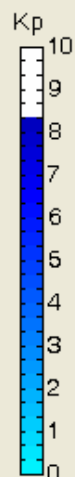
Dia

Hora UT

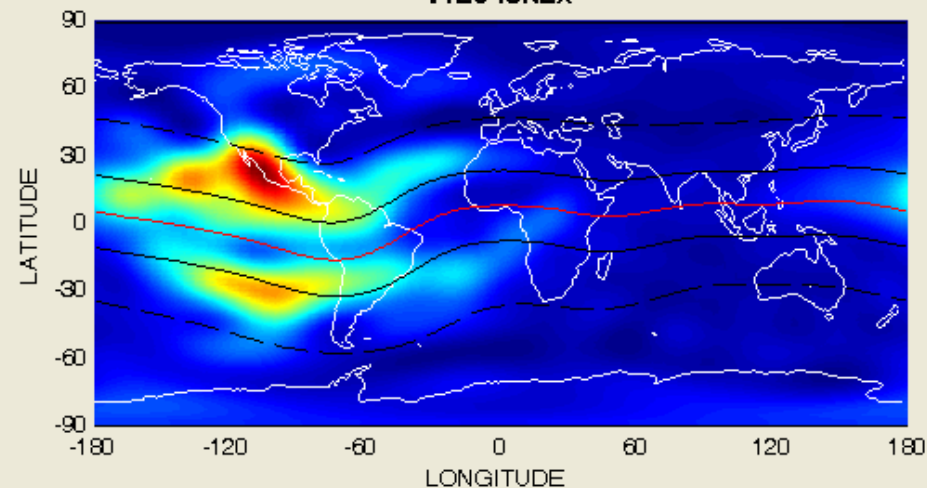
Plotar dados IONEX

Plotar VTEC do Traçado de raios

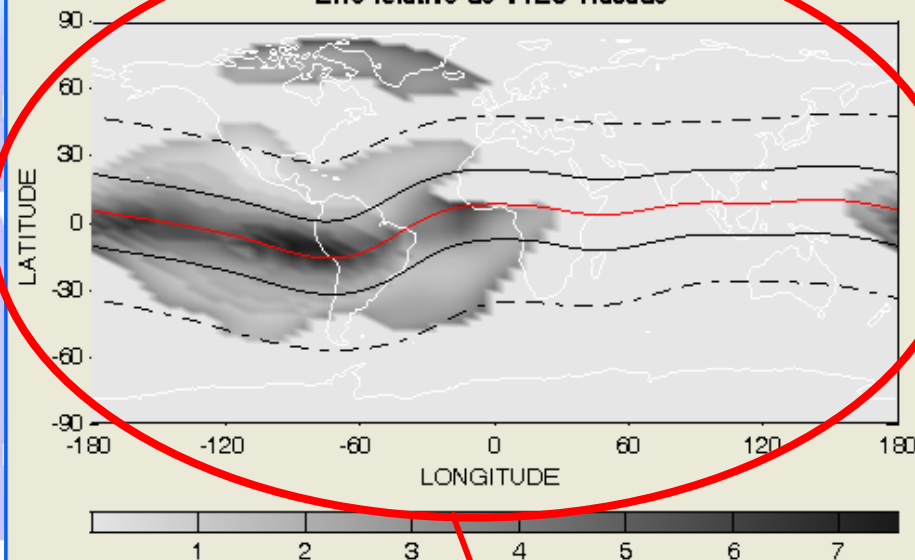
Plotar Diferença



VTEC IONEX

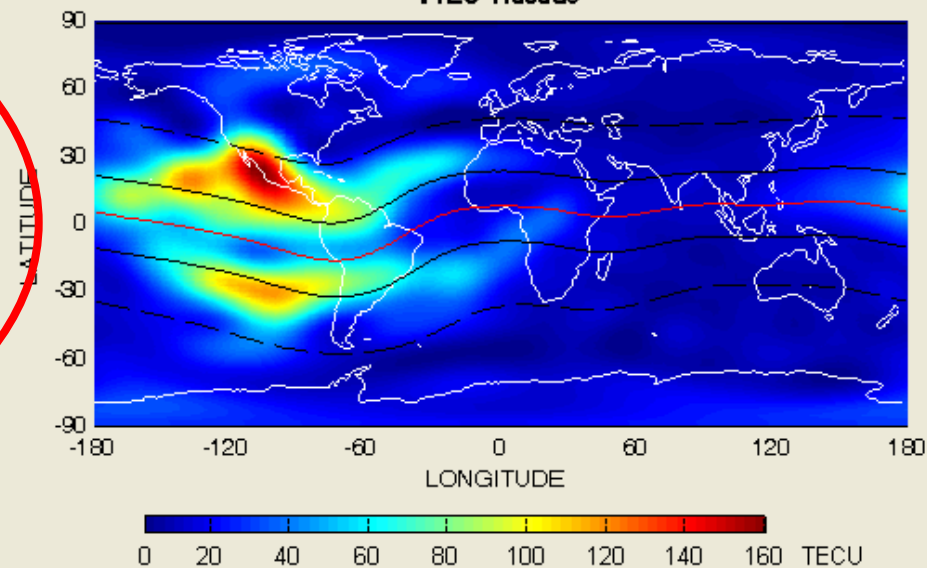


Erro relativo do VTEC Tracado



Relative error: $\times 10^{-3} (\%)$
< 15x10-3 %

VTEC Tracado



Sair



Acknowledgments

Part of this work was supported by FINEP (Financiadora de Estudos e Projetos), Grant nº 0/1/04/0441/00.

Thanks the organizers for sponsoring my participation in the URSI GA here in New Delhi as YS.

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THANK YOU
FOR YOUR ATTENTION

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THE END

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GLOBAL NORTH SOUTH STORY



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