



A new model for forecasting the Sub-Auroral Polarization Stream (SAPS).

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1 Extended Abstract

The Sub-Auroral Polarization Stream (SAPS) is a region of westward ionospheric flows equatorward of the main auroral oval which plays an important role in the mid-latitude space weather dynamics [1]. For example, SAPS has a controlling influence on the evolution of large-scale plasma features, such as storm enhanced density (SED) plumes, which in turn generate smaller scale irregularities that produce scintillation [2, 3]. This study examines the occurrence statistics of SAPS using a database of measurements obtained by the US mid-latitude chain of SuperDARN radars between January 2011 and December 2014. SAPS events were found to occur over a wide range of geomagnetic activity levels specified by the Dst-index, including non-storm and storm-time conditions. Spatial maps of SAPS occurrence probability and the most likely SAPS speed are presented for different Dst-index bins. During relatively quiet conditions ($-10\text{nT} < \text{Dst-index} < 10\text{nT}$) SAPS were observed 15% of the time and tend to be localized to the midnight sector and centered above 60° magnetic latitude. As the activity level increases, and Dst-index decreases, SAPS shifts equatorward and rotates duskward. During moderate storm conditions ($-75\text{nT} < \text{Dst-index} < -50\text{nT}$) SAPS are observed 87% of the time and tend to be centered at 20 magnetic local time and below 60° magnetic latitude. These results have been codified into a new empirical model which uses the Dst-index as input to estimate the probability of SAPS occurrence and most likely SAPS speed at a given magnetic latitude and local time. The accuracy of the model has been cross-validated using independent DMSP driftmeter data. Such a model will prove useful as a space weather forecasting tool as well as aiding future theoretical studies on SAPS formation and the generation of SAPS-associated small-scale ionospheric irregularities.

References

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